

3D geolocation of ionospheric plasma irregularities by combination of RO, in situ, and ground-based GNSS measurements



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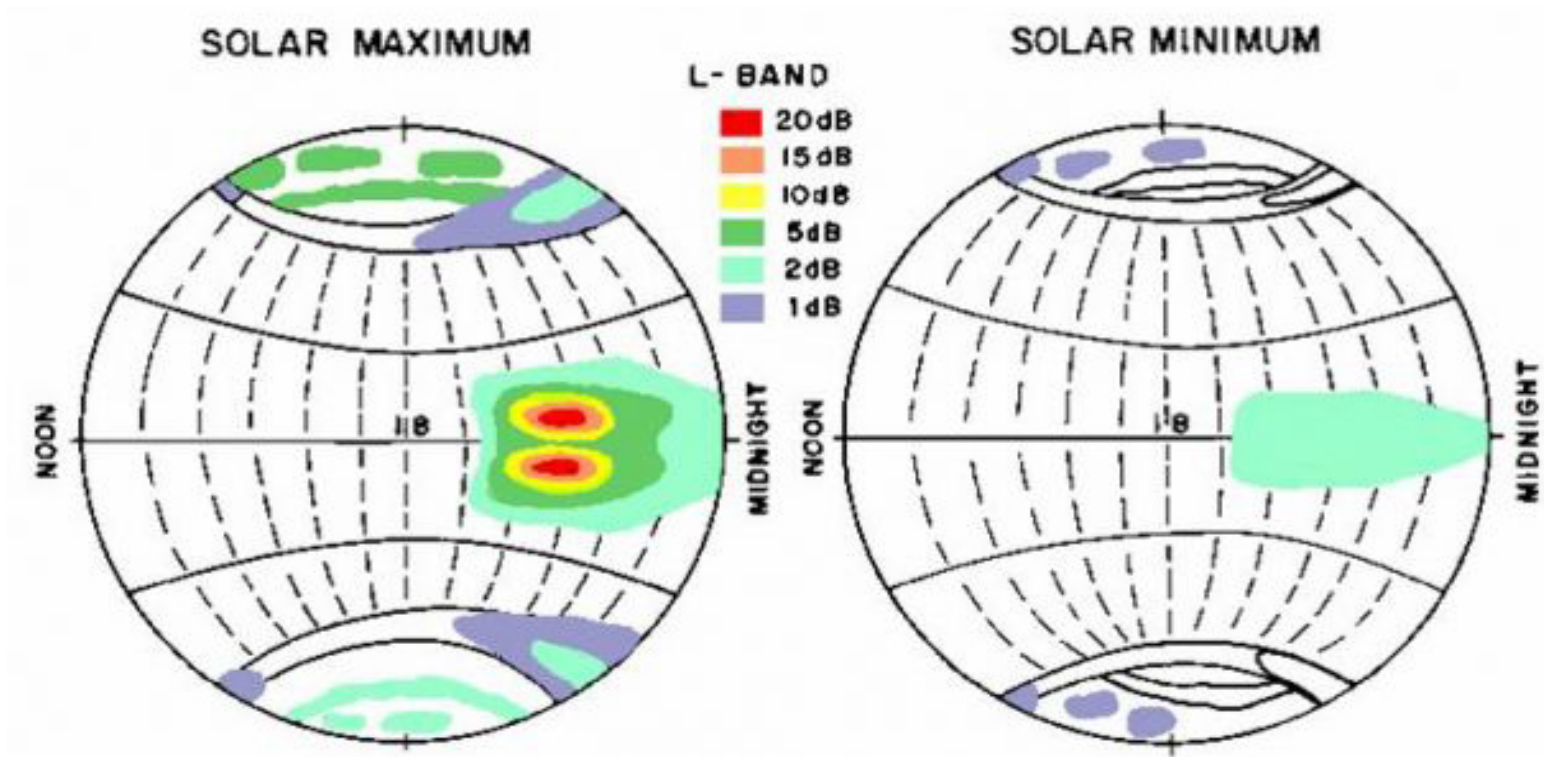


UNIWERSYTET
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W OLSZTYNIE

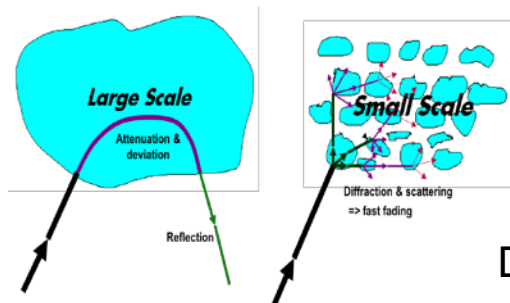
Irina Zakharenkova
SRRC UWM

Ionospheric plasma irregularities. Global distribution

Occurrence of L band scintillation during high and low solar activity

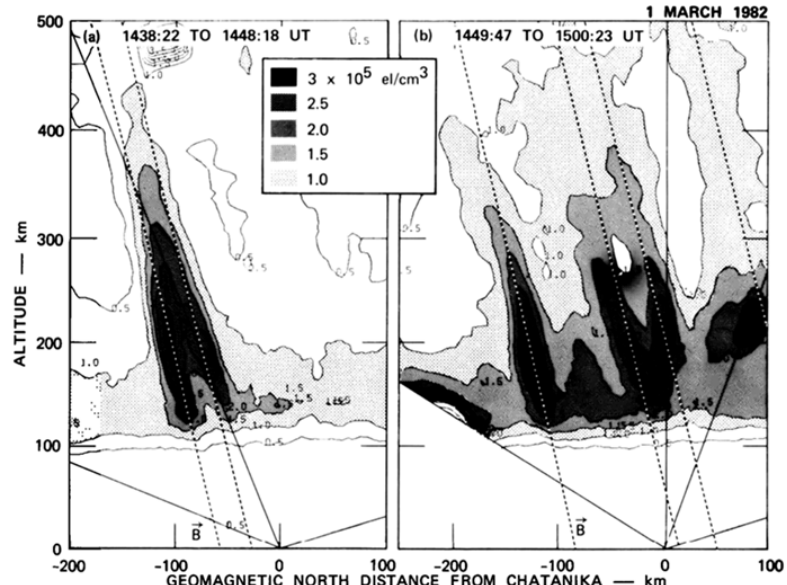


(Basu. et al., J. Atmos. Terr. Phys, 2002)

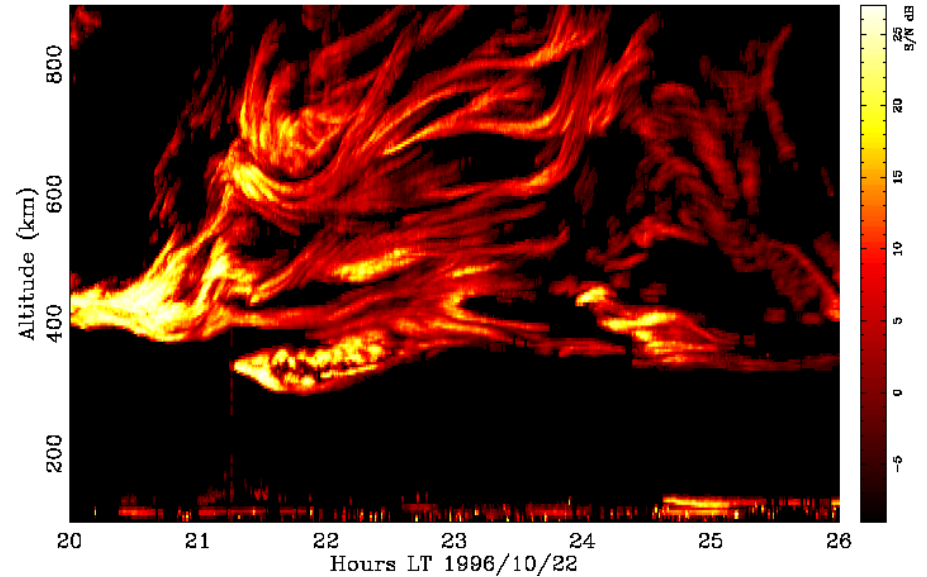


Different physical drivers leading to the structures generation

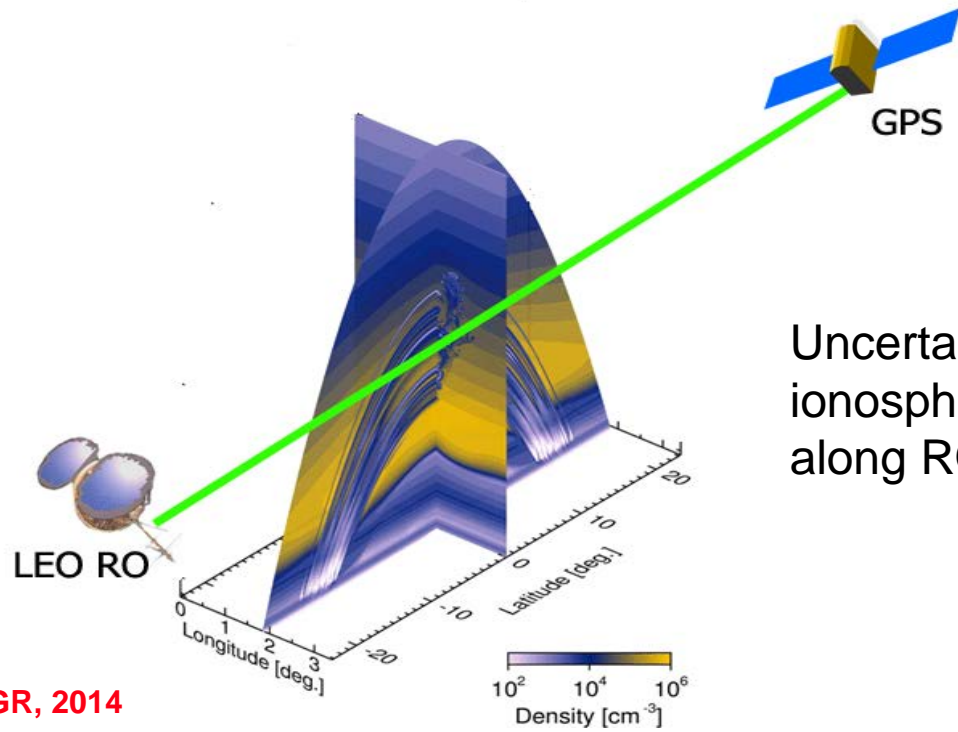
Spatial structure of ionospheric plasma irregularities



High-latitude irregularities over Chatanika
incoherent scatter radar

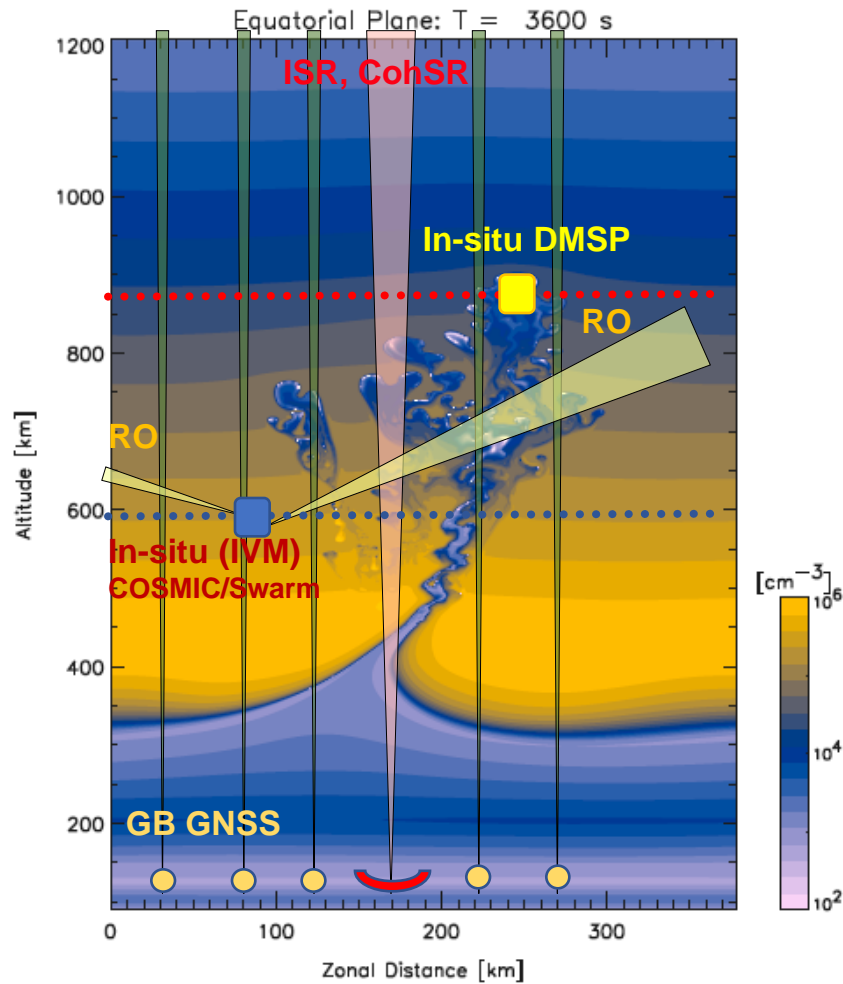


Altitudinal extent of equatorial irregularities
observed by Jicamarca radar



Uncertainties to locate
ionospheric plasma irregularities
along RO line of sight

Mutiinstrumental approach to specify 3d structure of ionospheric plasma irregularities



Challenge - source of irregularities / scintillations location and size estimation by single instrument

Using multiply instruments to detect plasma irregularities in 3D

Ionospheric irregularities signatures in GNSS measurements

Ionospheric irregularities can be characterized by measuring its impact on amplitude and phase of the received GPS signal.

Pi et al. [GRL, 1997] introduced into the use two GPS-based indices:

- **ROT** (Rate of TEC change, $dTEC/dt$) as a measure of GPS signal phase fluctuations

$$ROT = \frac{sTEC_k^i - sTEC_{k-1}^i}{t_k - t_{k-1}}$$

- **ROTI** (Rate of TEC Index, standard deviation of ROT) characterizes the severity of the GPS phase fluctuations

$$ROTI = \sqrt{\langle ROT^2 \rangle - \langle ROT \rangle^2}$$

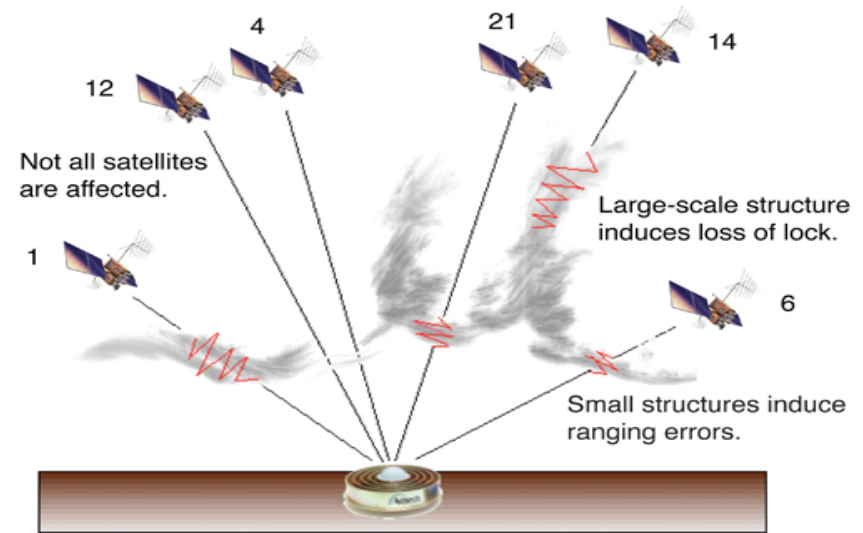
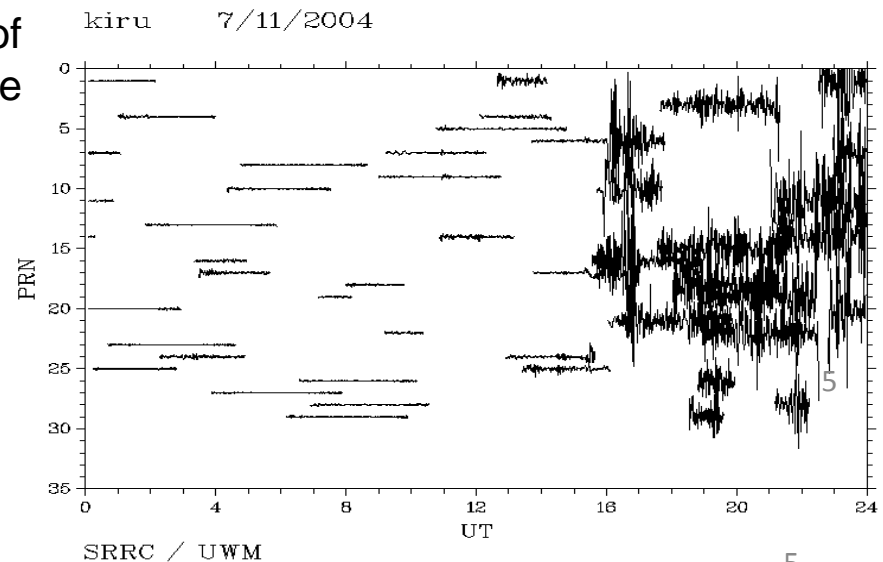
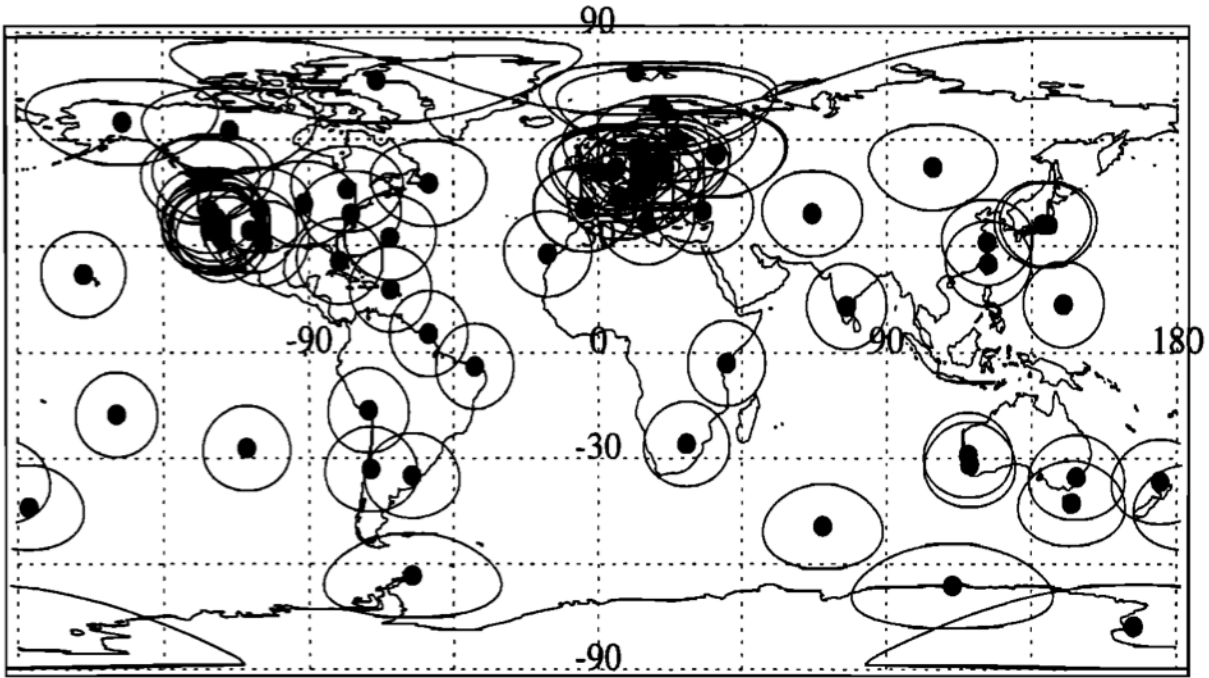


Image credit: GPS World

Example of ROT variations:



GNSS as global observational network for ionosphere monitoring

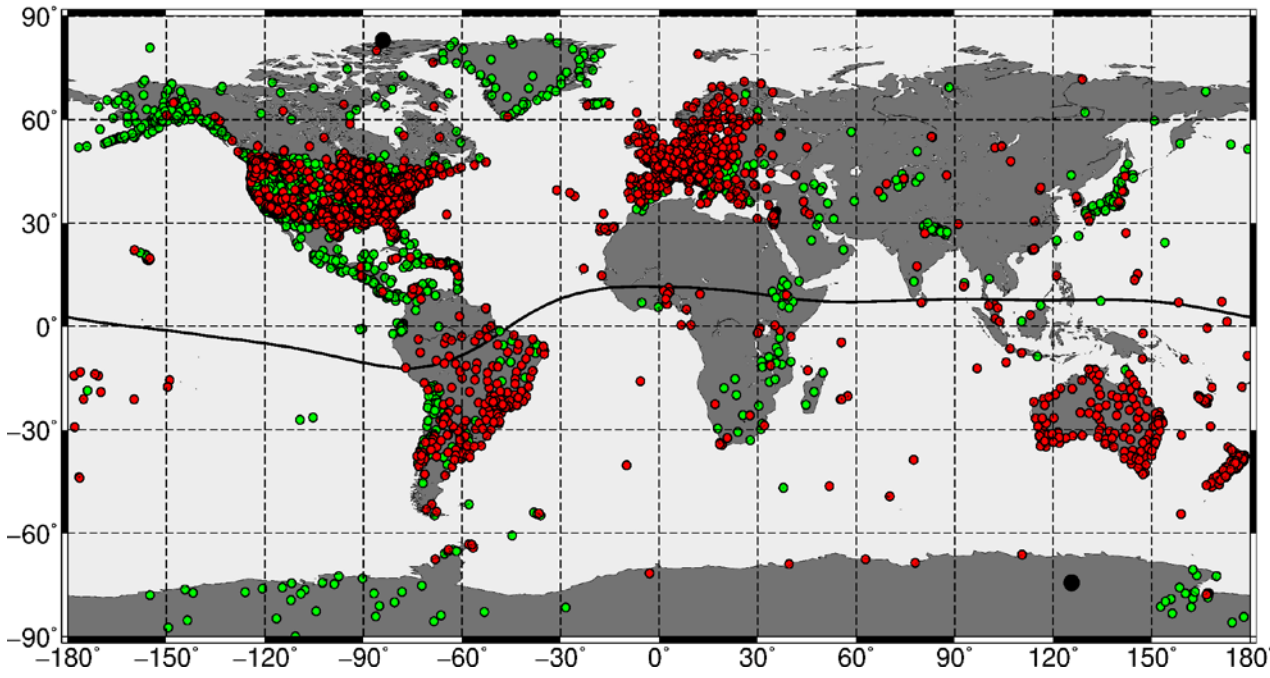


Year 1996
More than 100 GPS stations

Courtesy Manucci et al, 1998

GNSS Stations Distribution

- GPS
- GPS & GLONASS

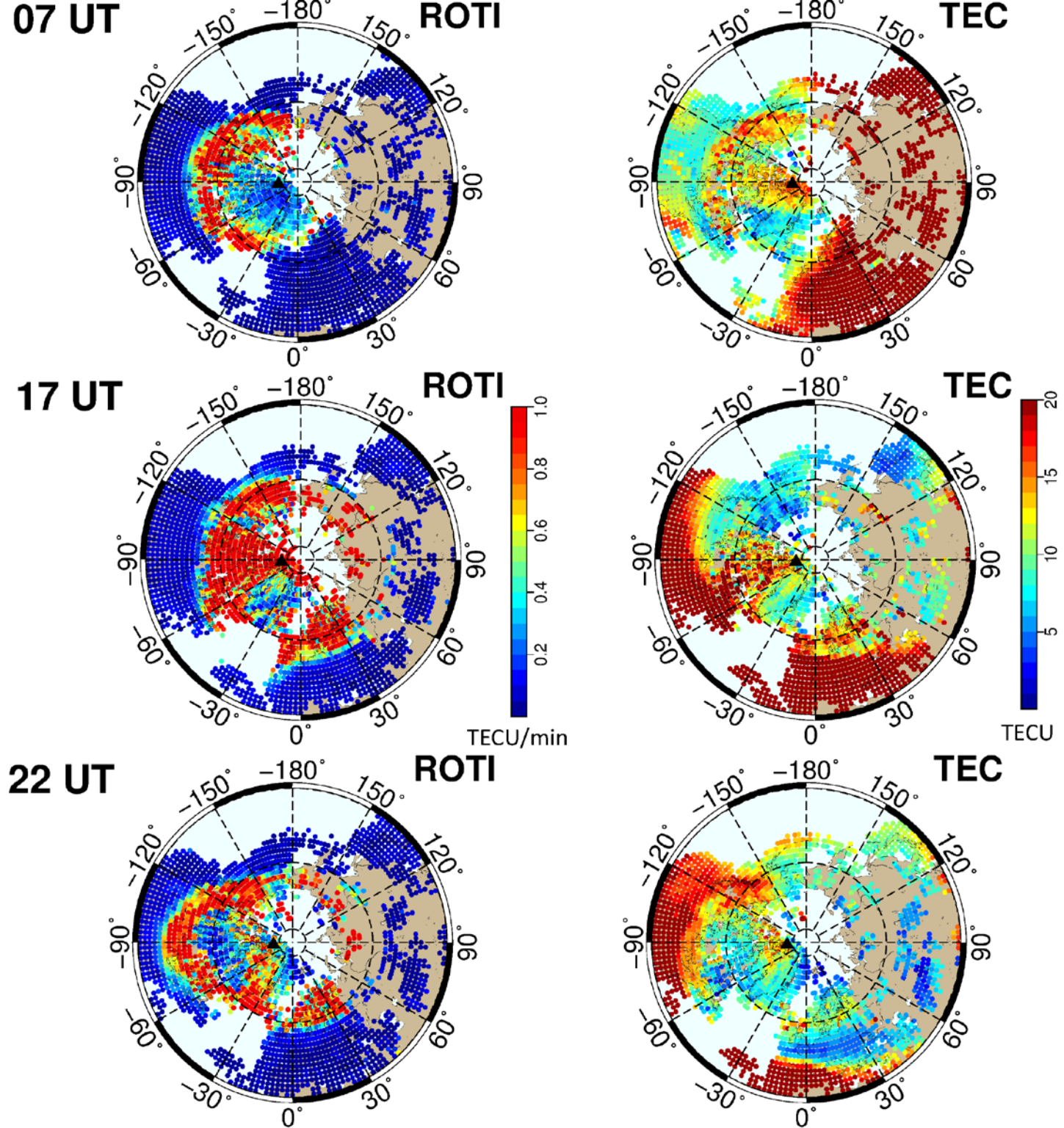


Year 2019
~ 6300 stations
~3500 multi-GNSS stations (GPS + GLONASS)

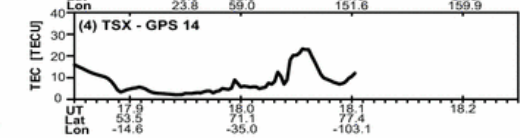
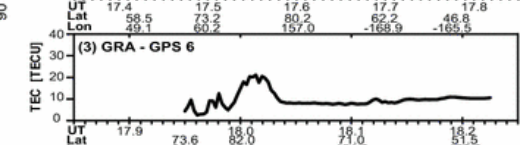
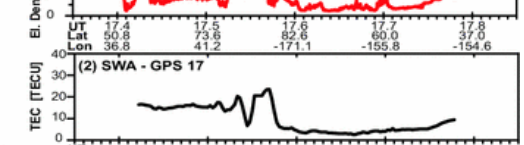
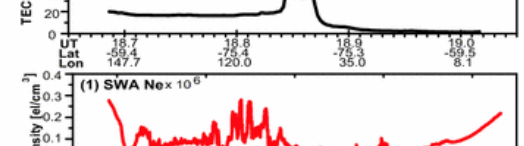
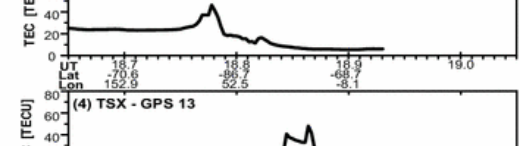
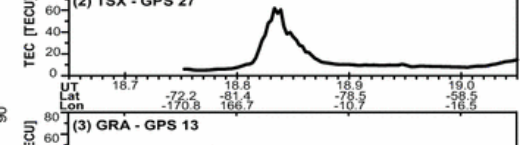
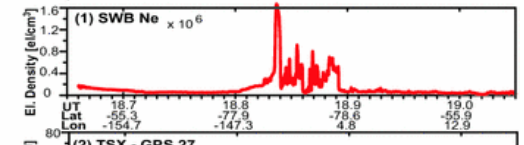
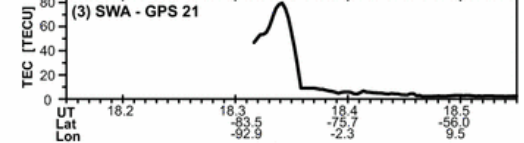
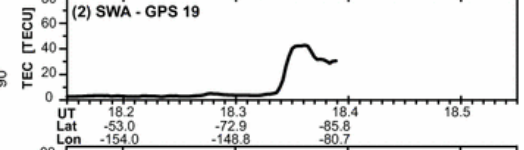
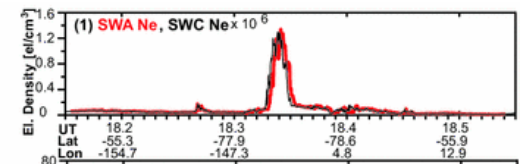
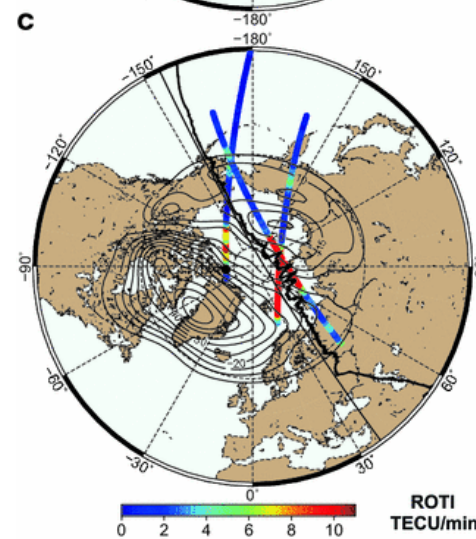
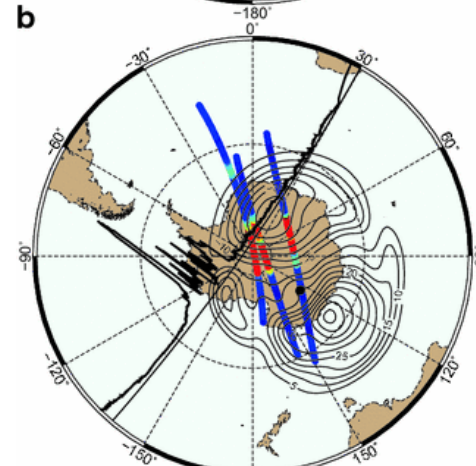
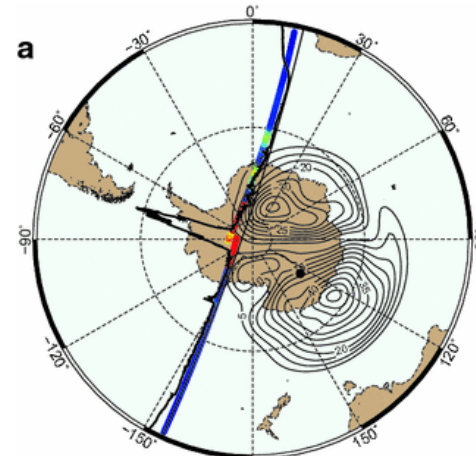
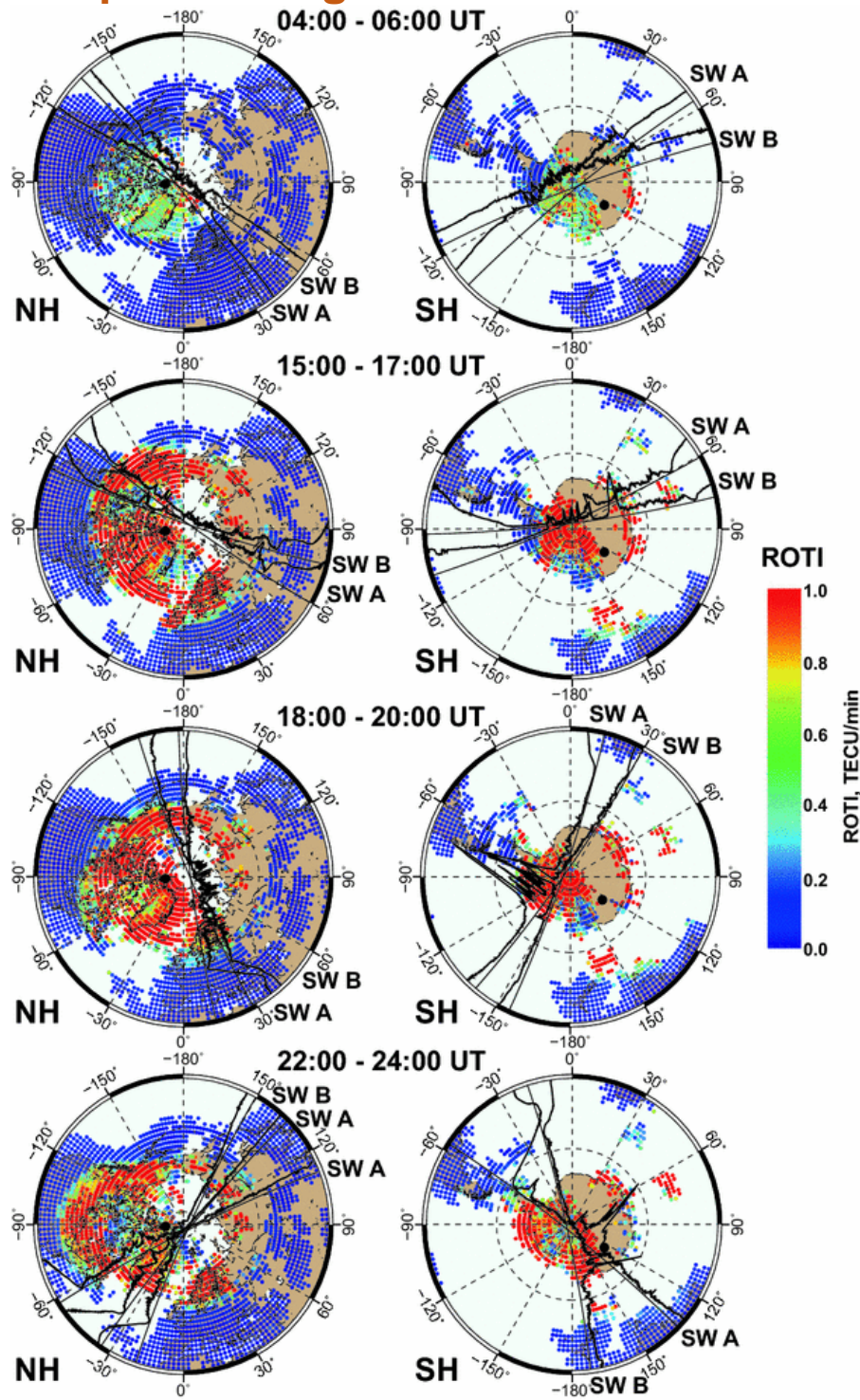
Ionospheric
plasma
irregularities
location.
SED-TOI structures
ROTI vs TEC maps

Ground based GNSS
observations

2015 St. Patrick's
Day storm



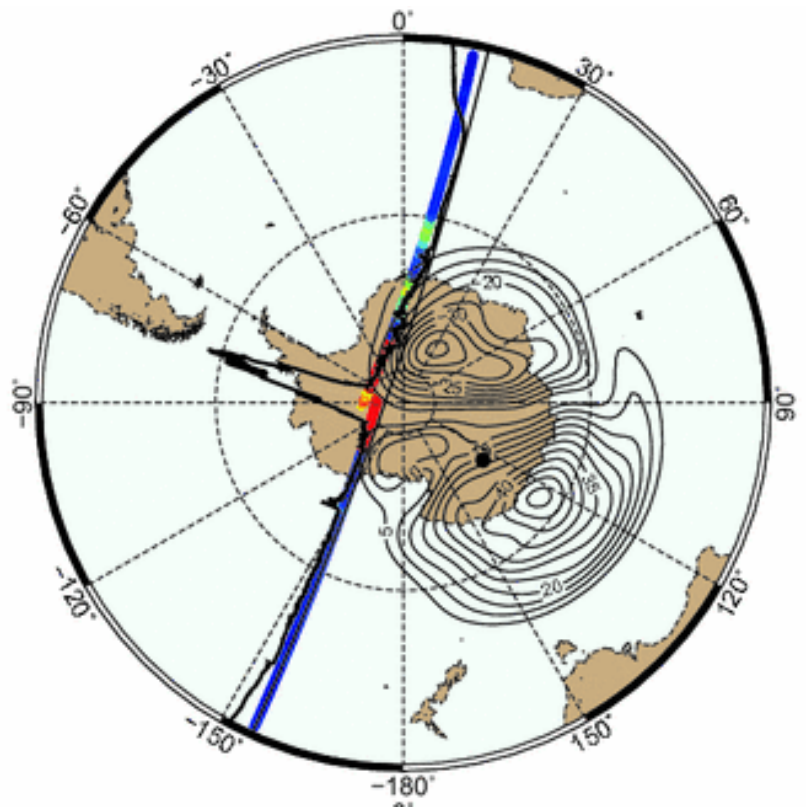
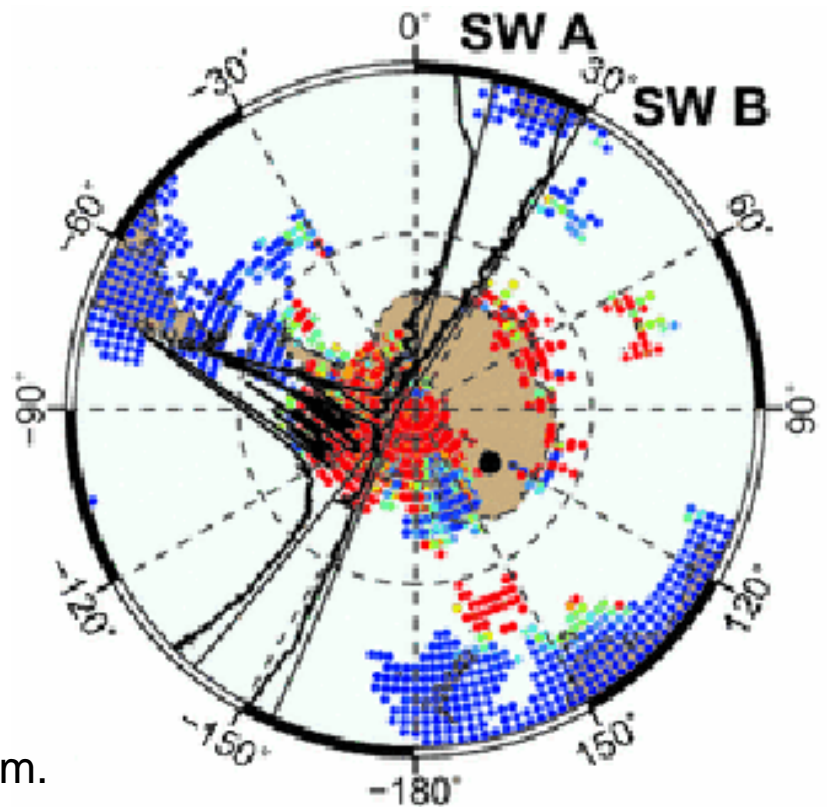
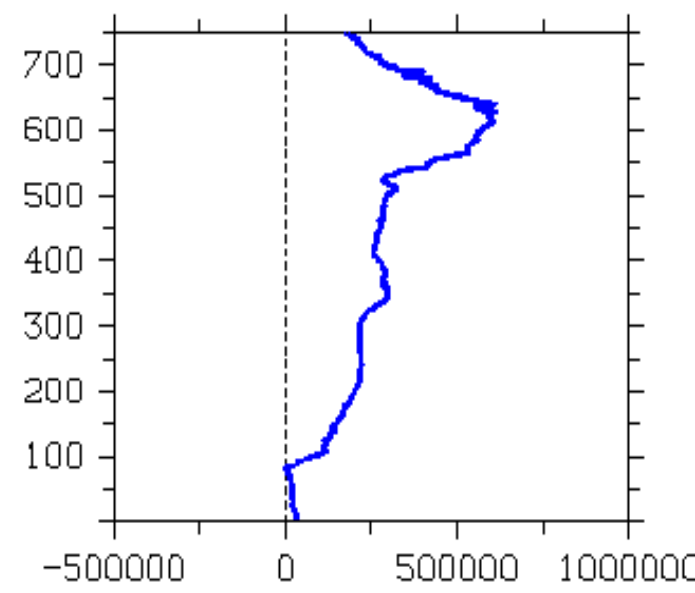
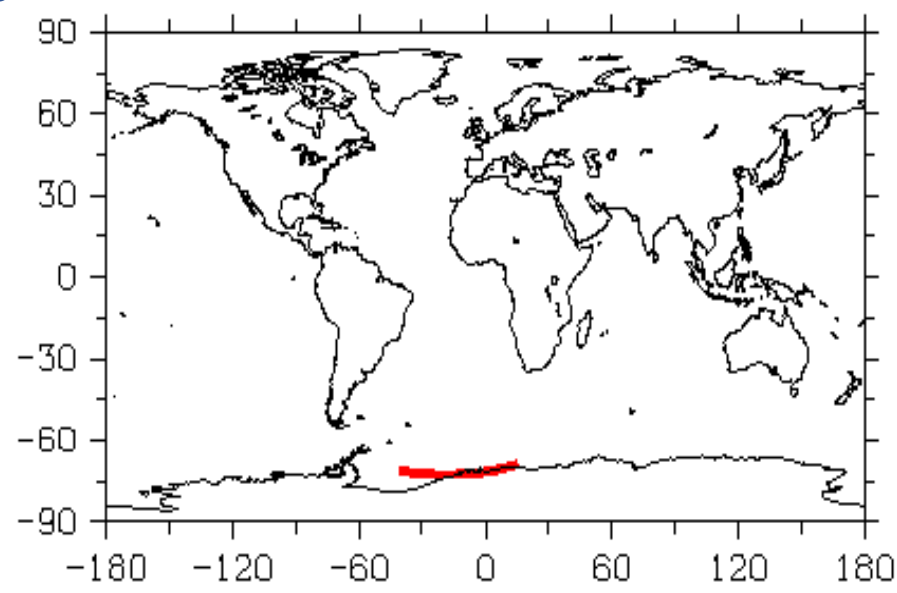
Ionospheric irregularities location. ROTI maps vs LEO GPS and in-situ observations



Ionospheric plasma irregularities location. SED/TOI signatures in the COSMIC RO

C006.2015.076.15.02.G16

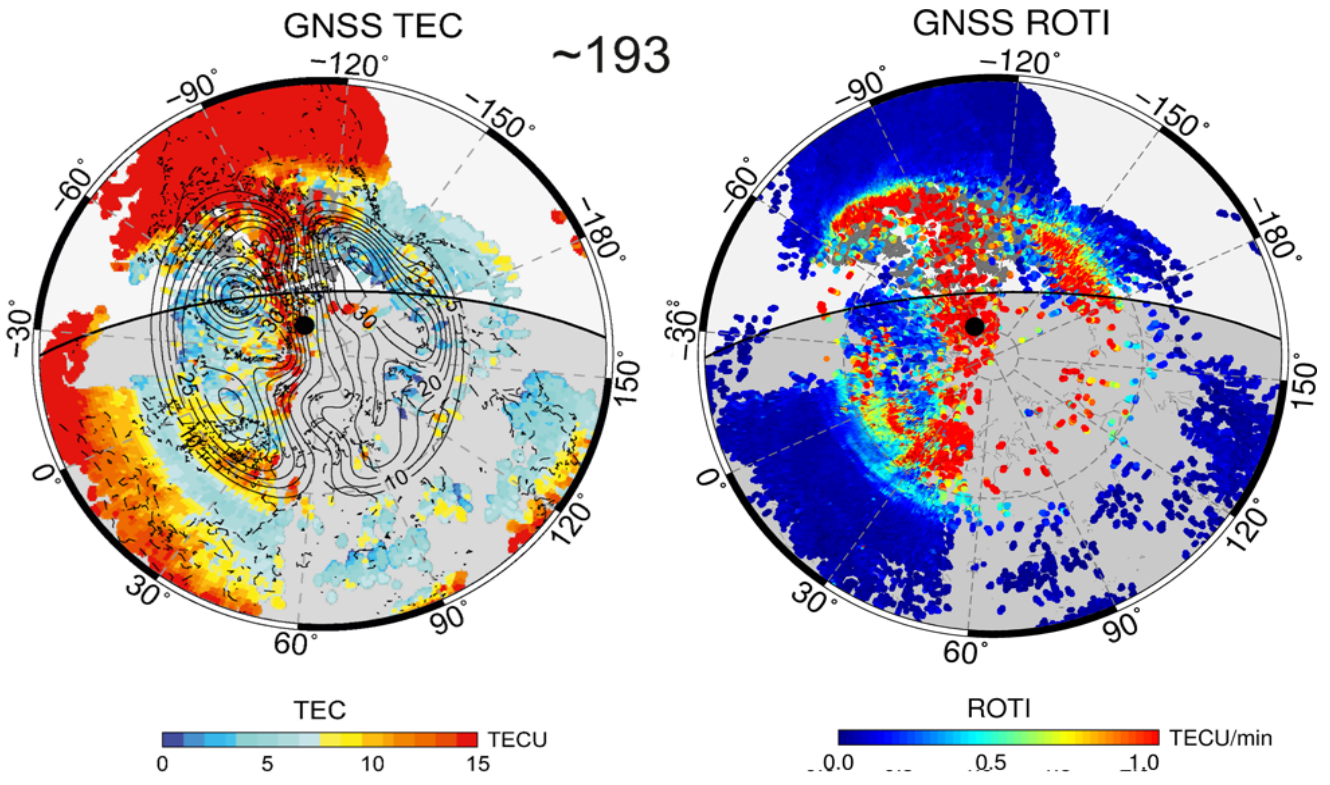
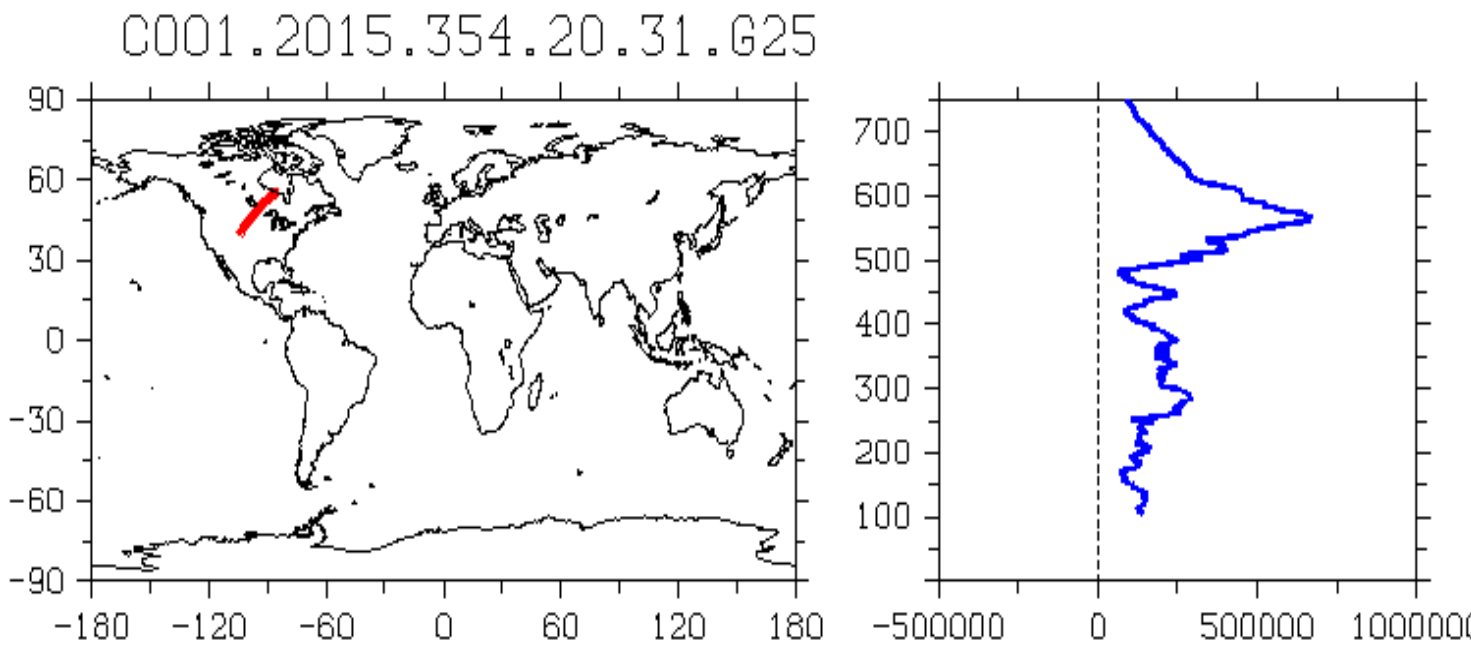
Superimposing to the
TEC/ROTI maps vs
LEO GPS and in-situ
observations



Formation of the
SED/TOI topside
structures over
Southern
hemisphere,
March, 2015
geomagnetic storm.

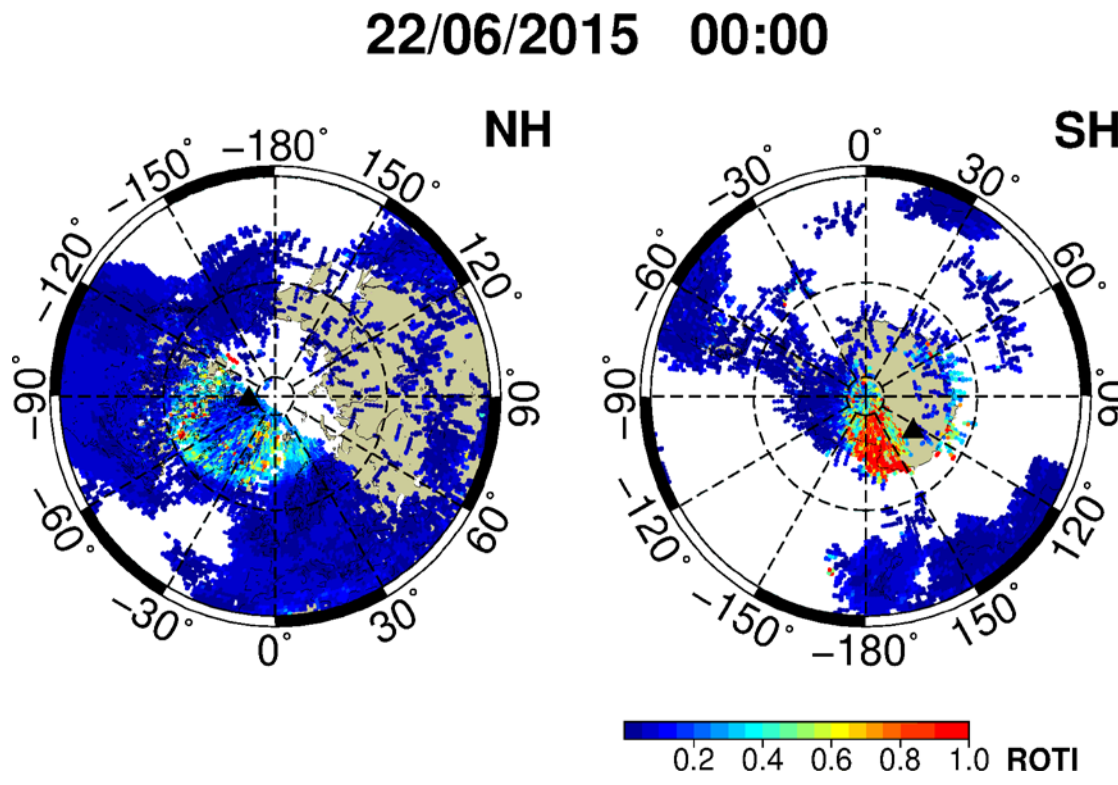
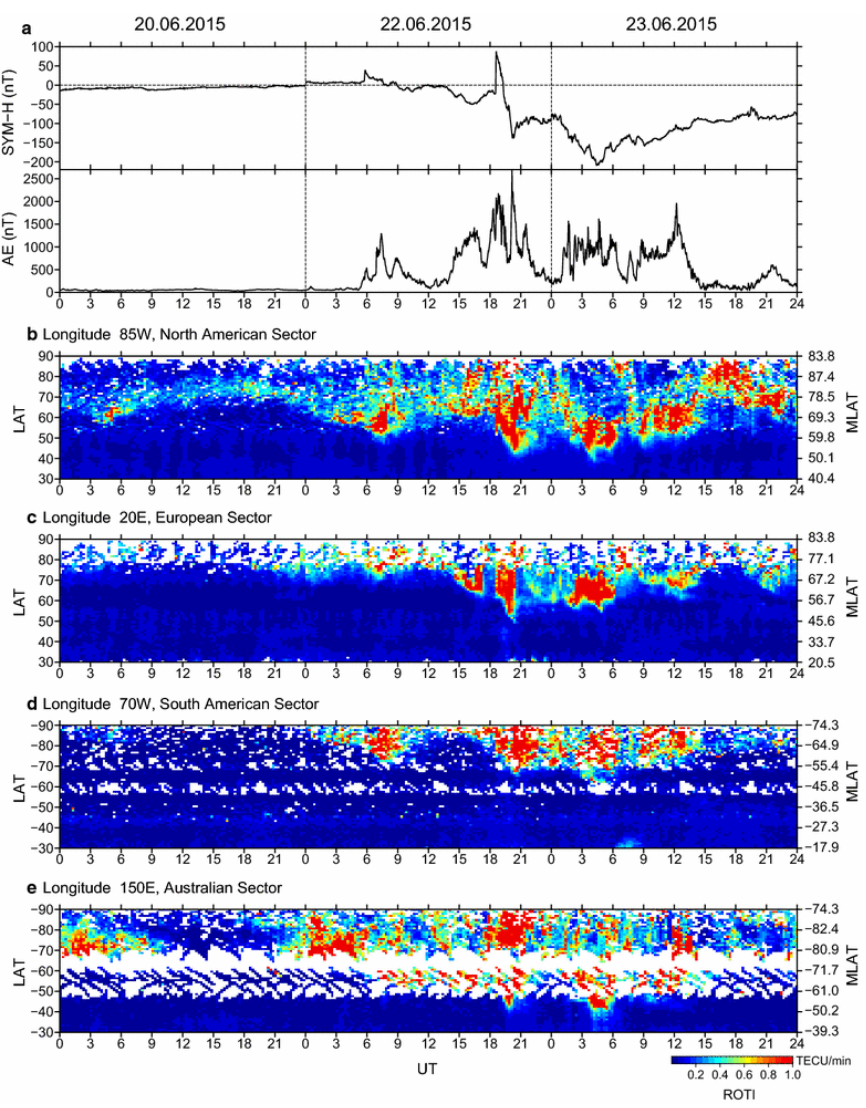
Ionospheric plasma irregularities location. SED/TOI signatures in the COSMIC RO

Superimposing to the TEC/ROTI maps vs LEO GPS and in-situ observations

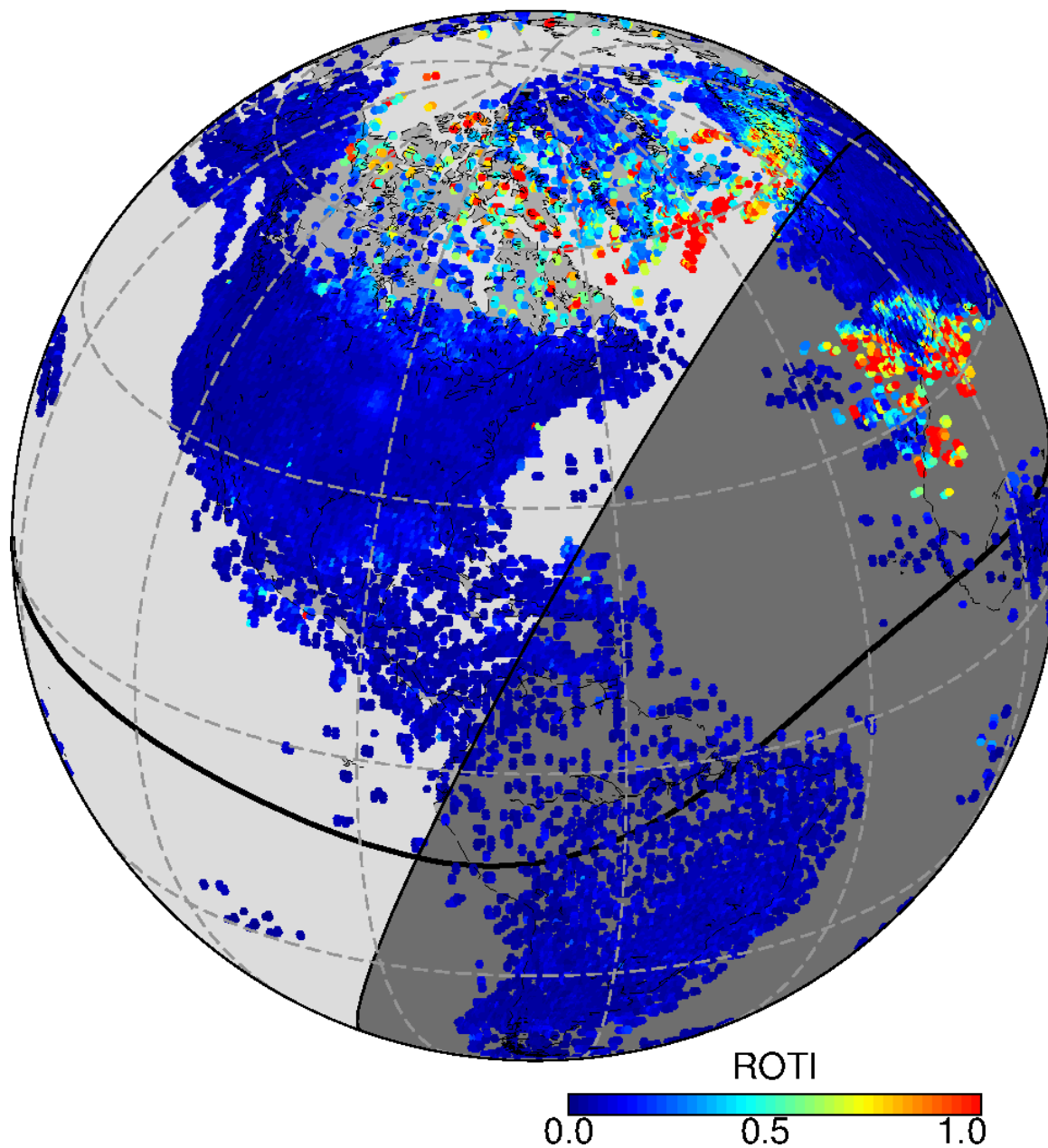


Formation of the SED/TOI topside structures over Northern hemisphere, December, 2015 geomagnetic storm.

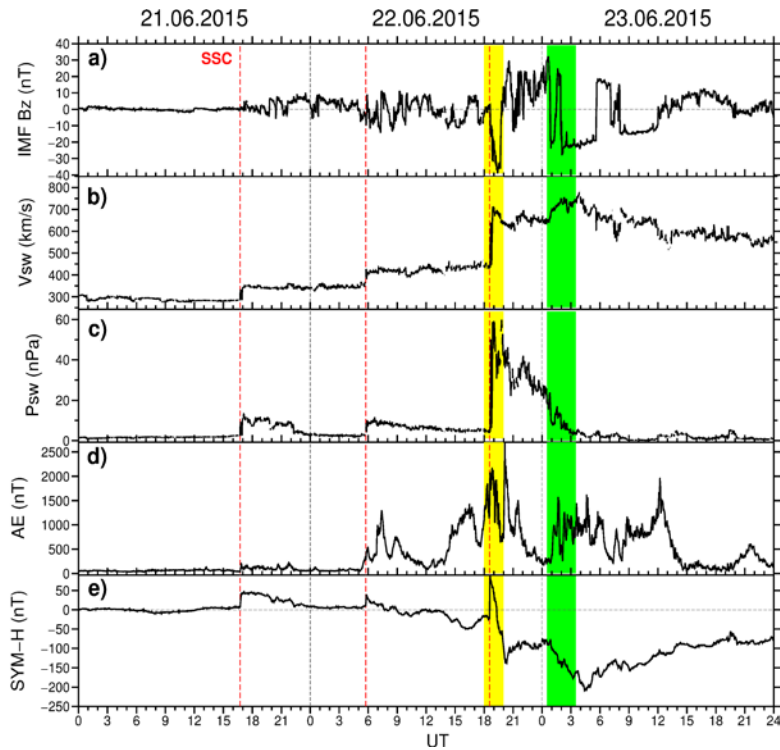
High-latitude ionospheric irregularities location vs Space Weather drivers



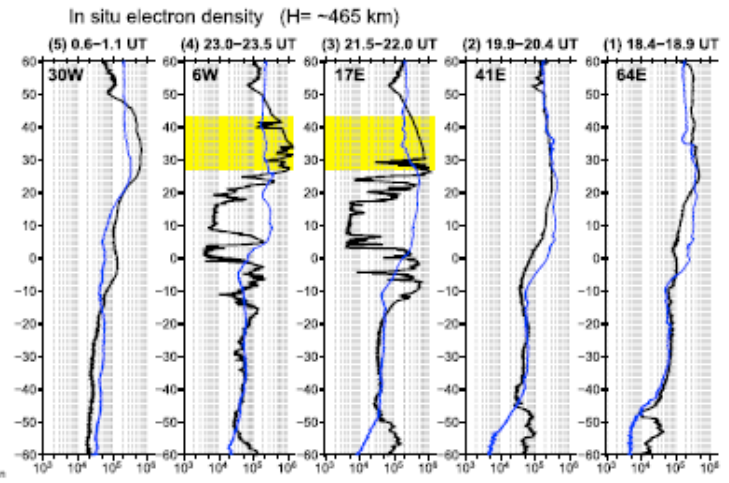
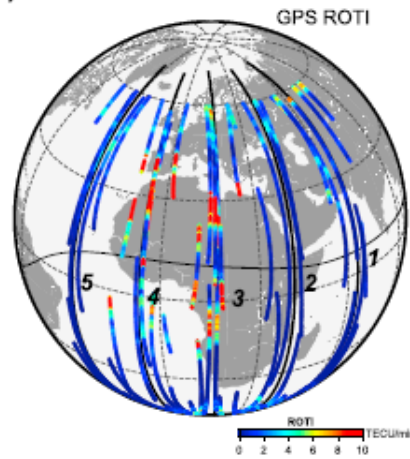
23/06/2015 0000 UT



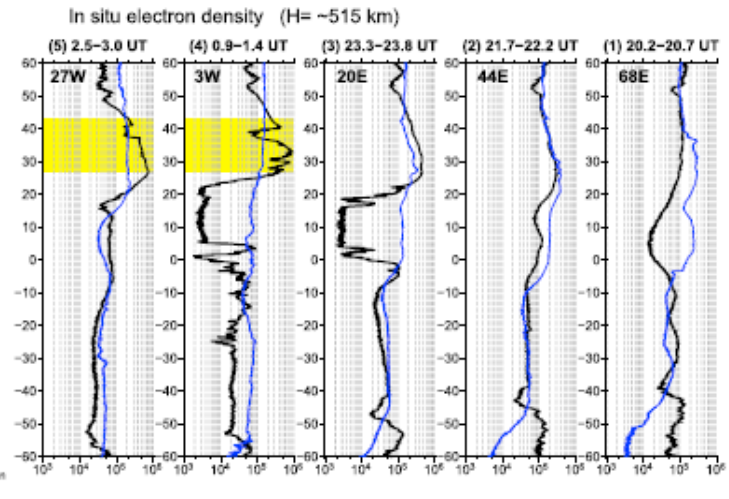
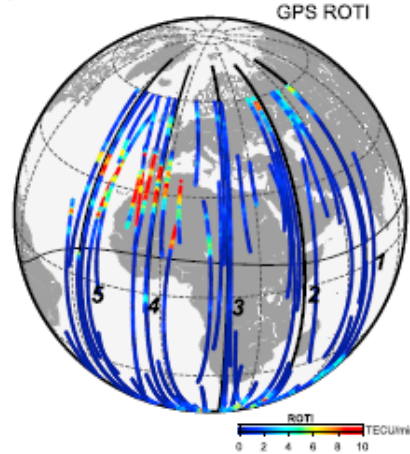
The June 2015 geomagnetic storm Plasma bubbles on midlatitudes



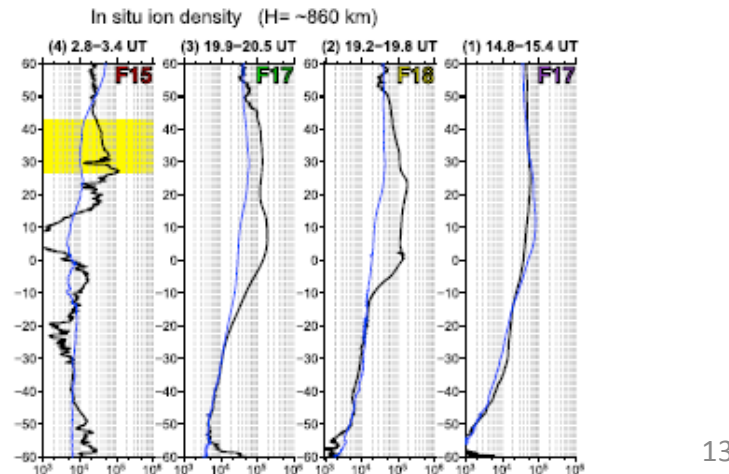
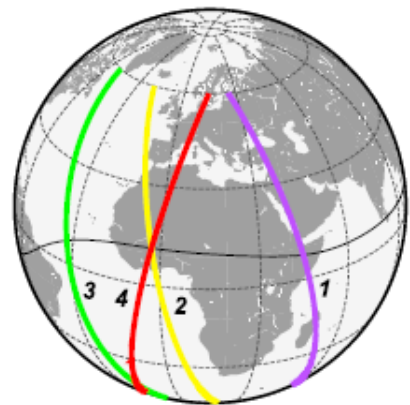
a) Swarm A



b) Swarm B



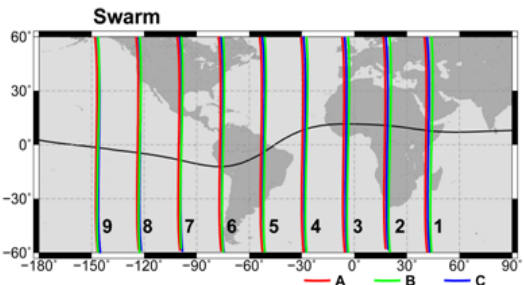
c) DMSP



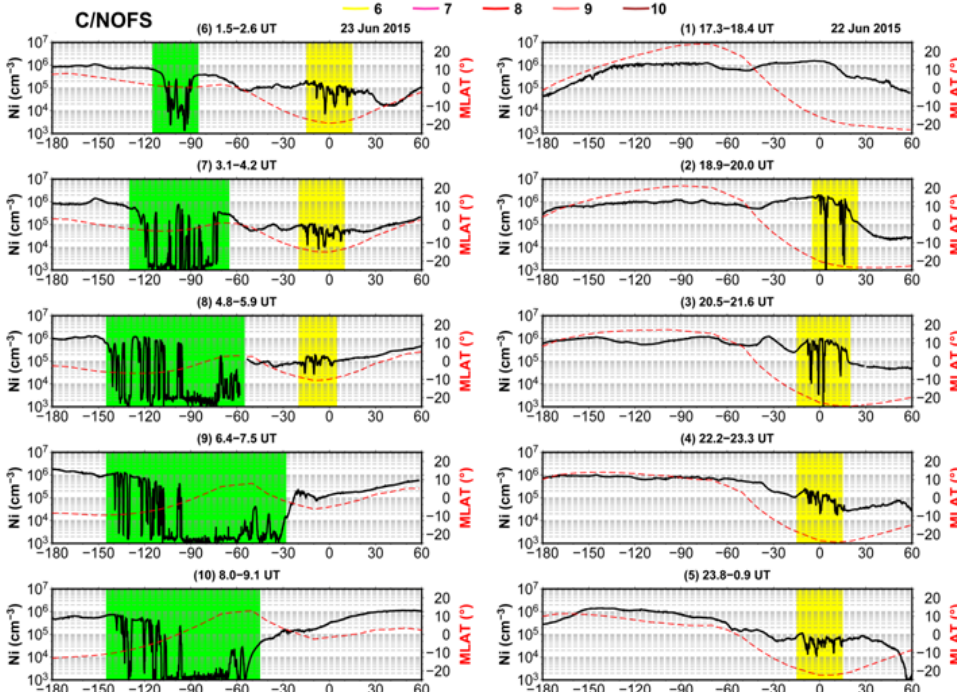
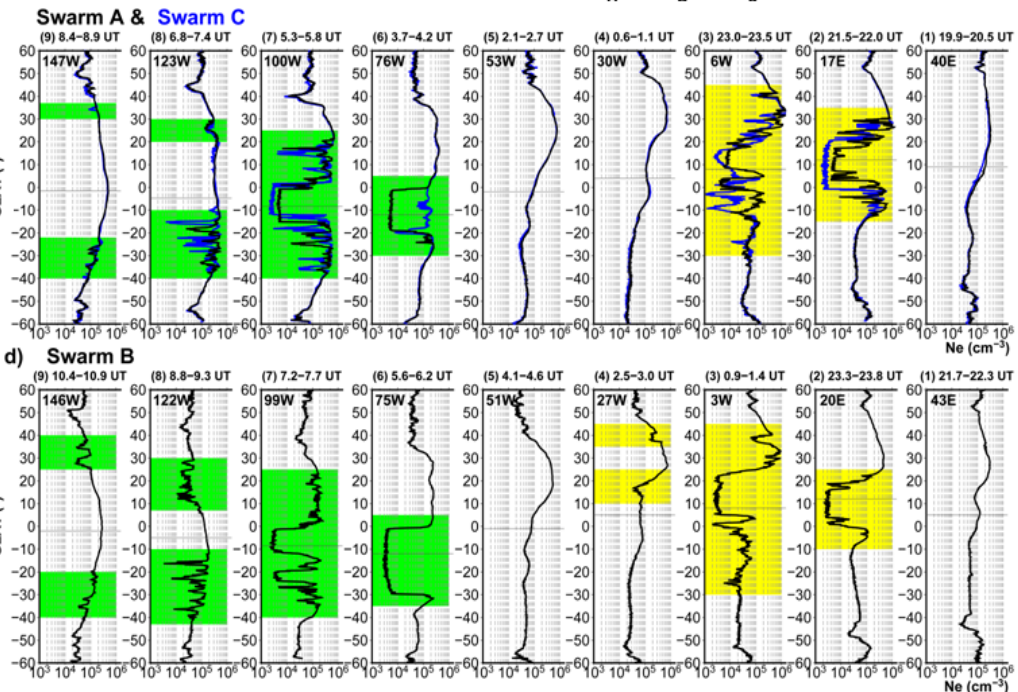
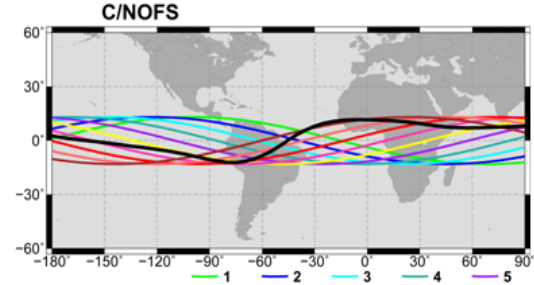
The plasma bubbles in signatures space-borne GPS ROTI and Ne from Swarm satellites passes and DMSP in situ ion density.

Swarm and C/NOFS in situ

polar orbits

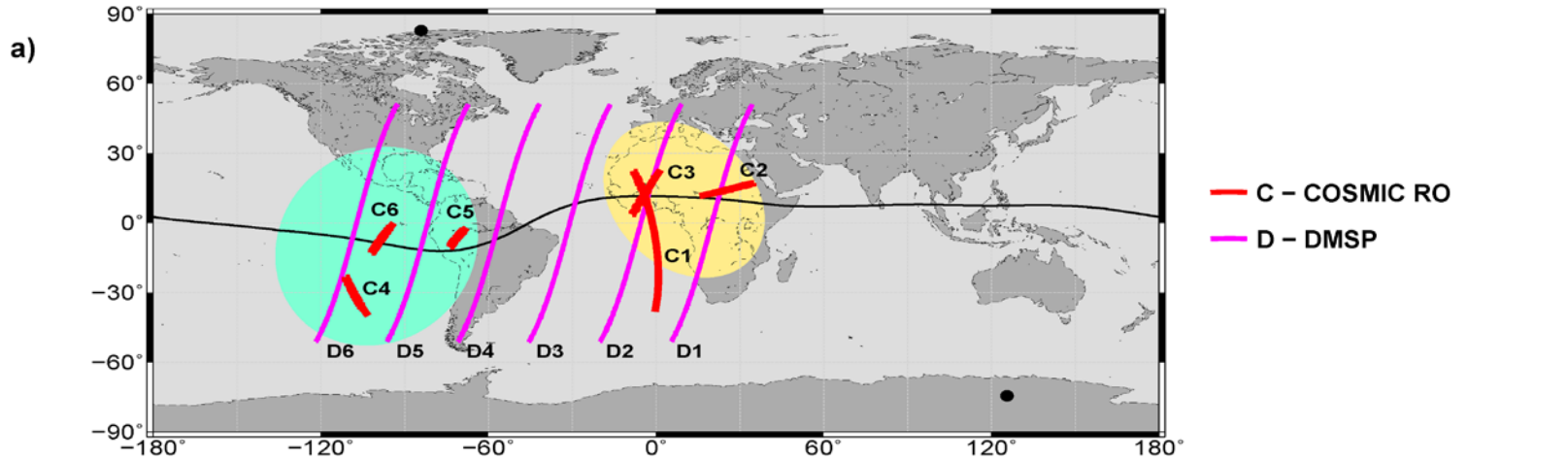


equatorial orbits

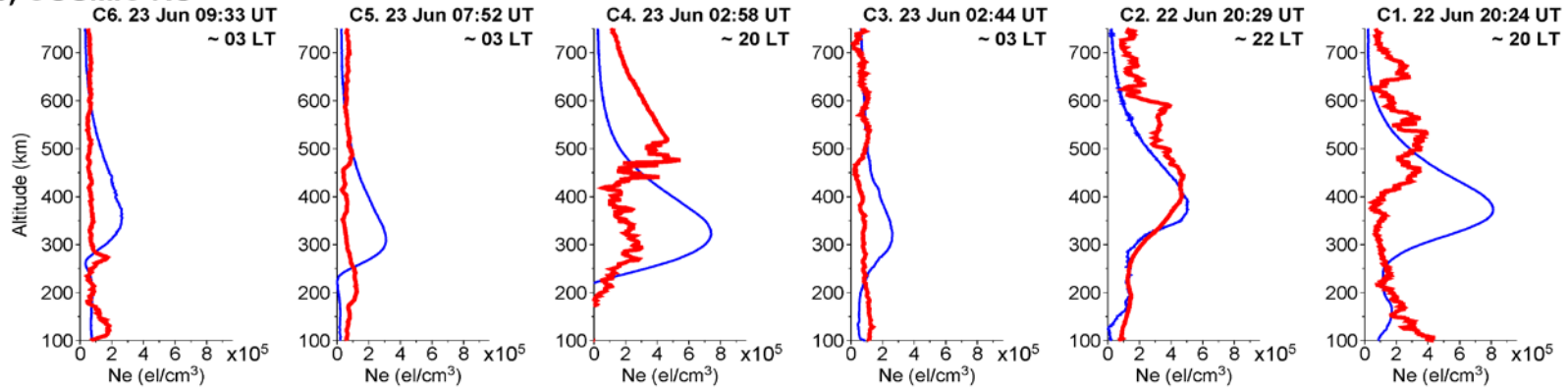


The June 2015 geomagnetic storm

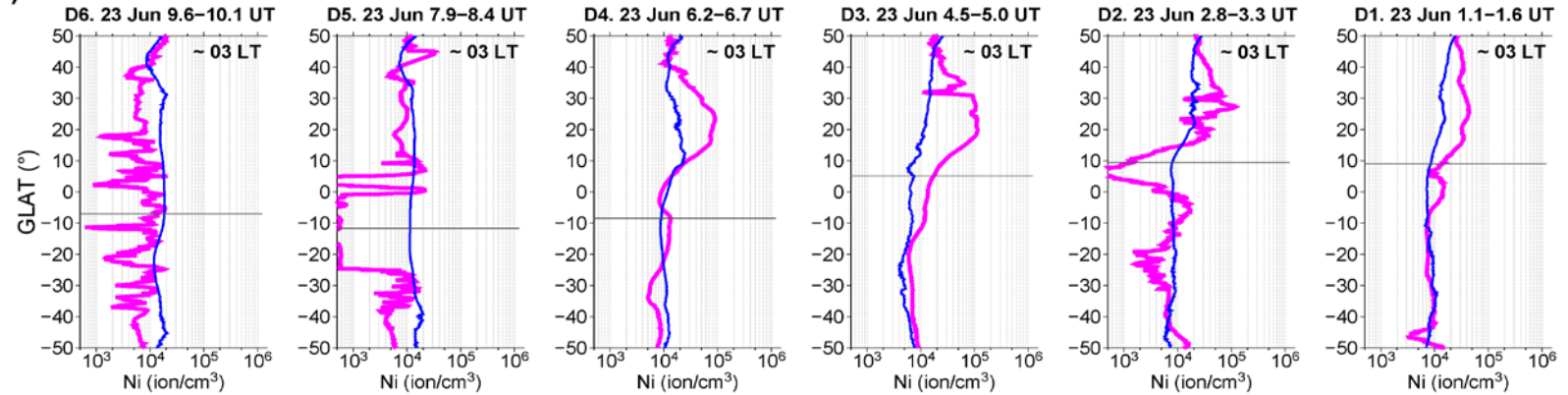
COSMIC RO vs in situ DMSP



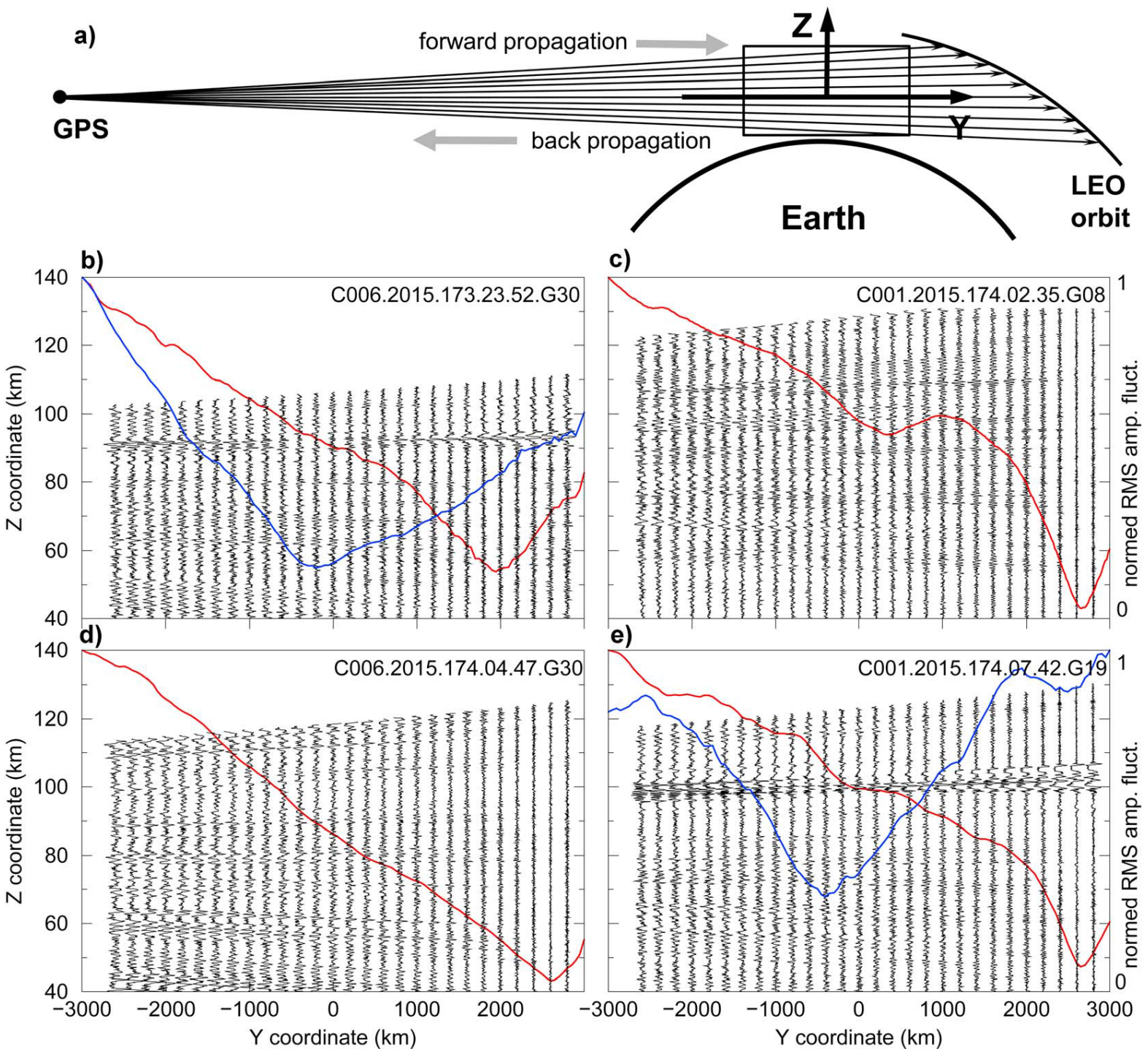
b) COSMIC RO



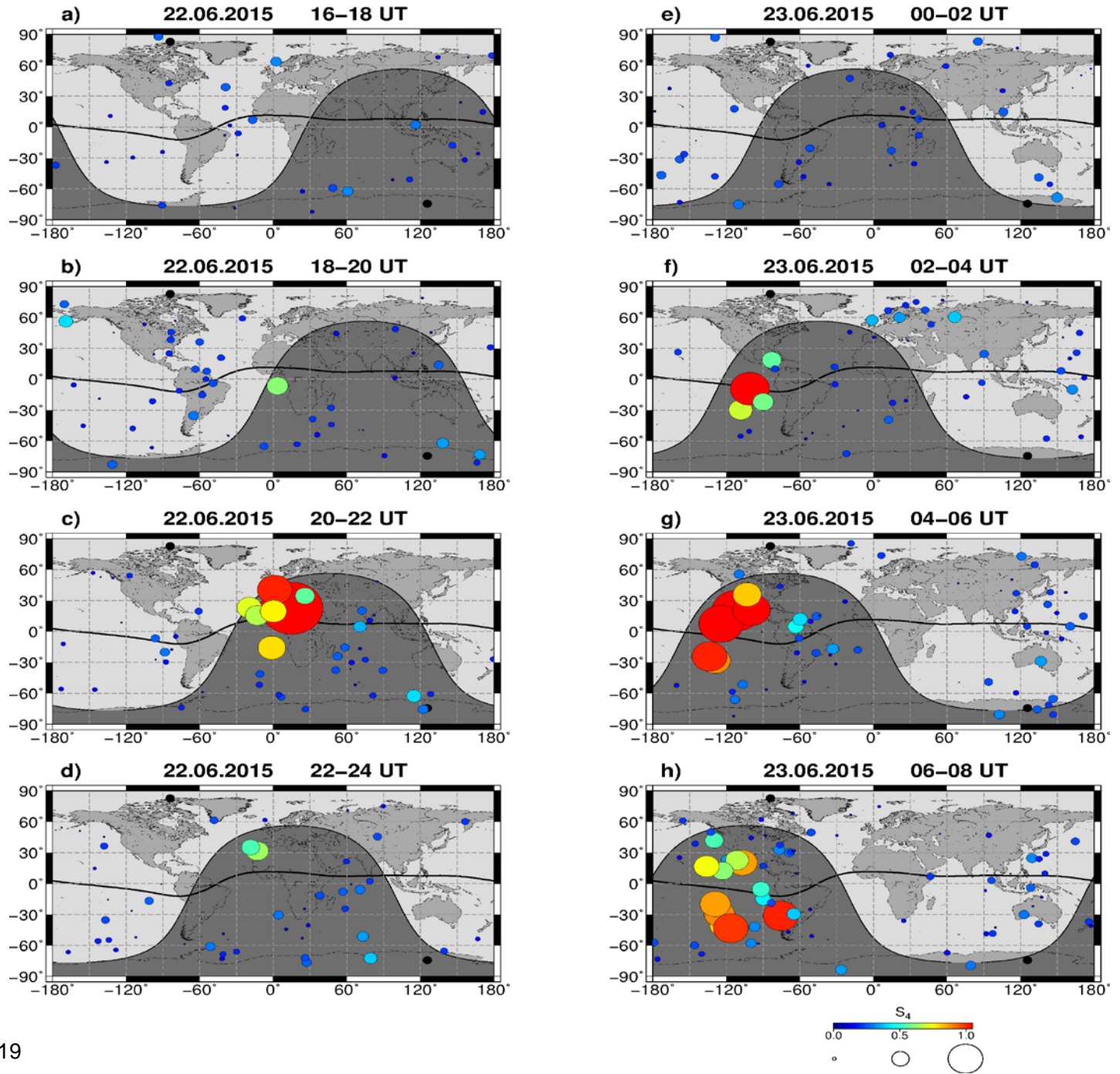
c) DMSP F15



Ionospheric plasma irregularities location by back propagation approach



COSMIC RO S4



Ionospheric plasma irregularities location by multi-instrumental observations

Measurements technique	Area of application
GNSS RO	Altitudinal distribution
RO BP	Distance from receiver, altitudinal distribution
GNSS POD	Location above satellite orbit
Ground based GNSS	Geographical location
In situ LP/IVM	Location along satellite orbit

Summary

- Combination of different ground-based and space-borne observations can allow to assess plasma irregularities parameters in geographical and altitudinal domain.
- Independent multi-instrumental measurements provide a consistent global view on the ionospheric density irregularity distribution and dynamics and allow estimating space weather drivers for plasma irregularities generation.
- Results confirm a high potency of the COSMIC-2 mission for ionospheric irregularities monitoring as it's equipped by both in situ plasma probes and GNSS remote sensing payload, including RO and POD instruments.

Thank you!