

# Introduction of Climate and Atmospheric Composition Exploring Satellites (CACES) Mission

**Liu Congliang (1,3), Sun Yueqiang (1,3), Bai Weihua, Du Qifei (1,3), Gottfried Kirchengast (2,3), Veronika Proschek (2,3)**

(1) National Space Science Center (NSSC), Chinese Academy of Sciences, Beijing, China

(2) Wegener Center for Climate and Global Change (WEGC), University of Graz, Graz, Austria

(3) JOAC–Joint Laboratory on Occultations for Atmosphere and Climate by NSSC and University of Graz

**Thanks to all further colleagues at LSEE (RO Group) and IAP, CAS (CACES group) valuable discussions and advice!**

## Outline

- ◆ 1. Radio Occultation Group, NSSC
- ◆ 2. From GRO to LEO-LEO Occultation
- ◆ 3. CACES Mission
- ◆ 4. LMO Research in NSSC for CACES
- ◆ 5. Summary and Conclusion

## ➤ Research domains:

- GNSS radio occultation (GNSS RO)
- GNSS reflectometry (GNSS-R)
- **LEO-LEO microwave occultation (LMO)**

## ➤ Research Goals:

- To develop occultation and GNSS-R instruments
- To apply the GNSS remote sensing data

## ➤ Research partners:

- China Meteorological Administration (CMA)
- Institute of Atmosphere Physics (IAP), CAS
- Wegener Center for Climate and Global Change (WEGC)
- Geo Forschungs Zentrum (GFZ)
- DMI, EUMETSAT ROM SAF
- Institut d'Estudis Espacials de Catalunya (IEEC)



ISSI BJ Forum Group photo



Prof Kirchengast and IAP colleagues



GNSS RS Group and Prof Gottfried Kirchengast

## GNSS RO Receivers

GPS RO receivers

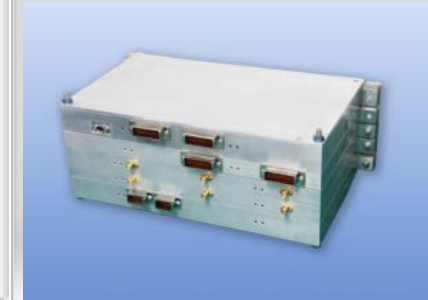


FY-3C/-3D  
GPS/BDS  
compatible  
flight models

FY-3C GNOS

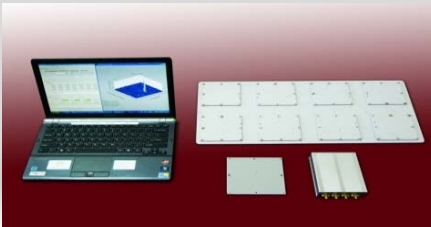


FY-3E GNOS II



## GNSS Reflection Receivers

GPS-R  
software  
receiver



FPGA+DSP  
GPS/BDS/GALILEO  
compatible ocean  
reflection receiver

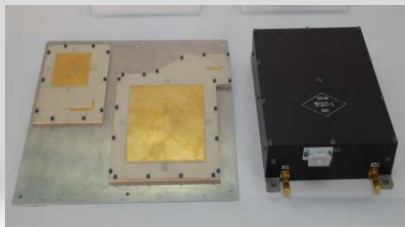


**LMO research fields:**

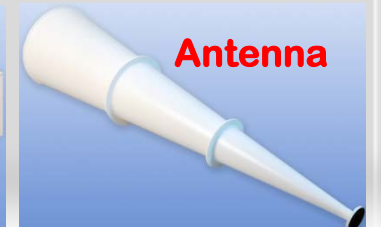
- Satellite-to-Satellite occultation for Mars
- LEO-LEO occultation for Earth

## LMO Transmitters and Receivers

YH-1 RO  
receiver  
flight model  
for Mars



LMO  
transmitter  
receiver  
for Earth





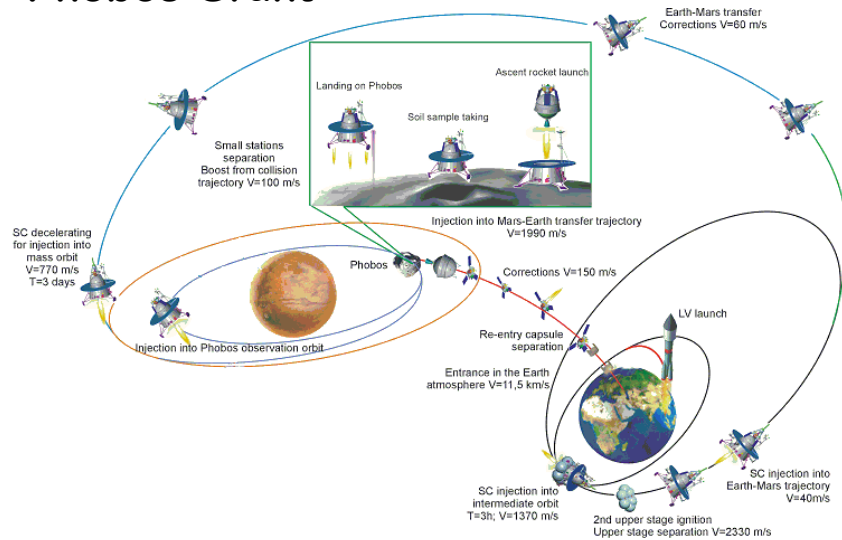
# YingHuo-1 mission



Phobos-Grunt



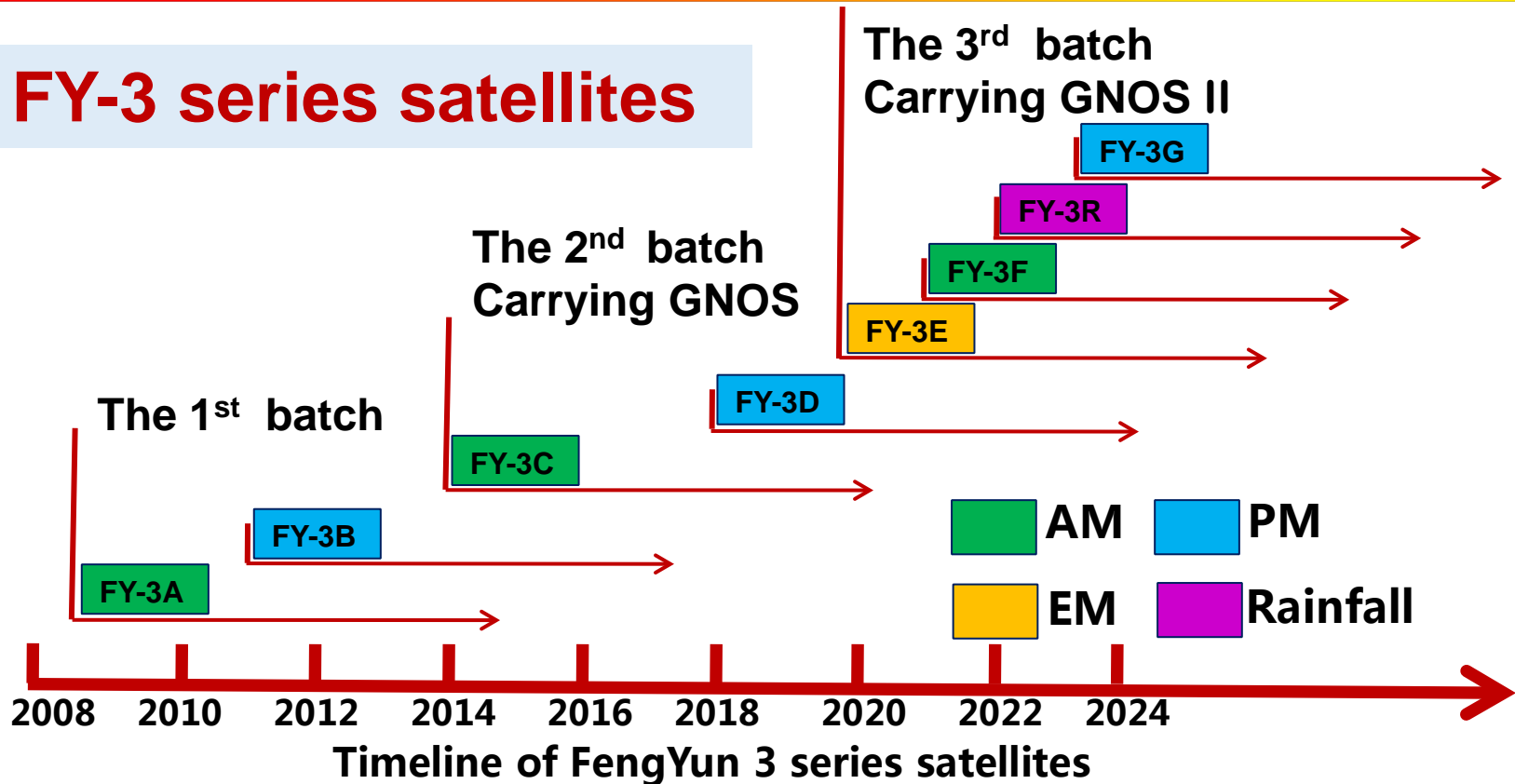
HY-1 micro-satellite structure



- One of the major scientific objectives of YH-1 mission is **to investigate the ionosphere of the Mars by using Sat-Sat occultation.** (The 2 frequencies signals are  $\sim 415.5$  and  $\sim 833$  MHz.)
- The indoor testing of the Tx and Rx showed that both of them were working well.
- However, they were not successfully launched into the Martian orbit in 2011
- And this is the start point of RO research group in NSSC.

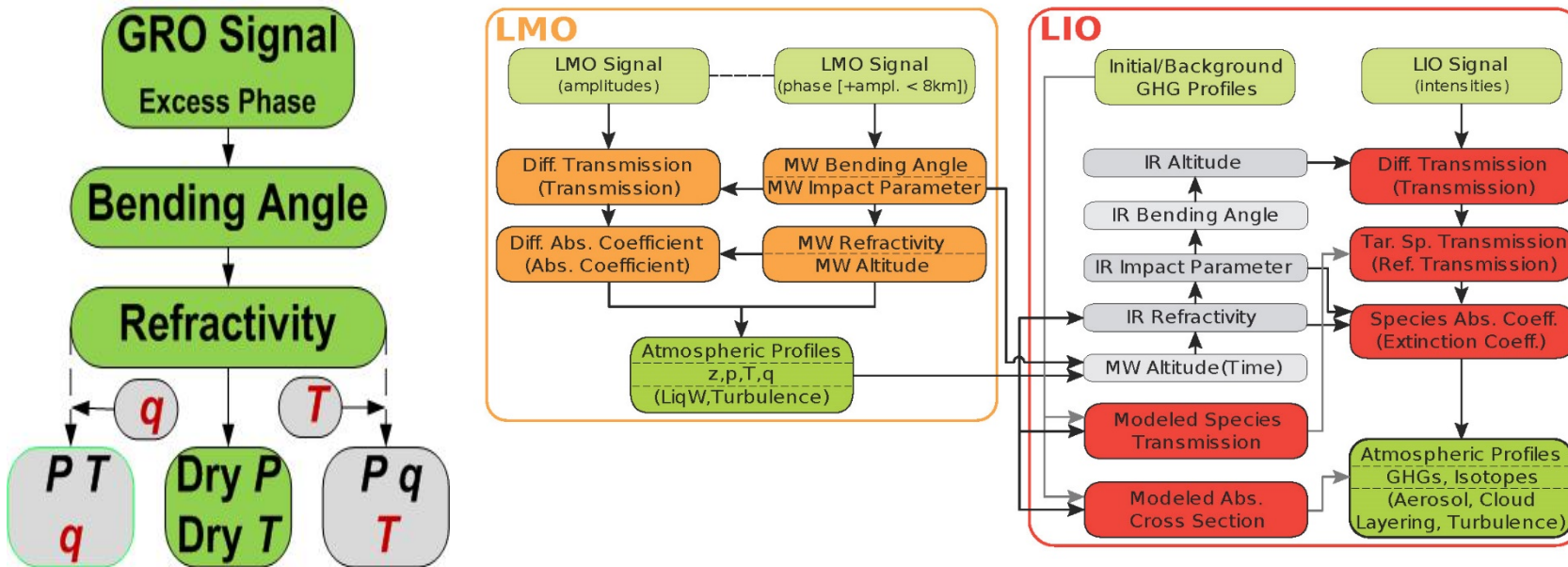


## FY-3 series satellites



- FY-3C GNOS was launched in September 2013
- FY-3D GNOS was launched in November 2017
- FY-3E GNOS II plan to be launched in 2020
- The following FY-3 satellites will carry GNOS II as a key payload...

## Retrieval processing of GRO, LMO and LIO (LMIO) :



[V. Proschek, G. Kirchengast, and S. Schweitzer, AMT, 4, 2035–2058, 2011]

GRO has been established by various missions (COSMIC , MetOP, FY-3 GNOS...)

Widely used for NWP, climate monitoring and space weather.

**Temperature-humidity ambiguity**, which means the temperature and humidity in the troposphere can only be retrieved separately from refractivity by co-using a priori humidity or temperature information

GRO cannot realize the atmosphere GHGs monitoring

The complex refractive index is given by:

$$n = n' + i \cdot n''$$

The optical path is given by:

$$L = \int_{Tx}^{Rx} n' ds$$

Excess phase
Bending angle

Abel transform to refractivity real part:

$$n'(a) = \exp \left[ \frac{1}{\pi} \int_a^{\infty} \frac{\alpha(a')}{\sqrt{a'^2 - a^2}} da' \right]$$

The optical depth is given by:

$$\tau = \int_{Tx}^{Rx} k ds$$

Amplitude

$T = I/I_0 = \zeta \exp(-\tau)$

Abel transform to absorption coefficient :

$$\Delta k(a) = -\frac{1}{\pi} \frac{da}{dr} \int_a^{\infty} \frac{d\Delta\tau / dx}{\sqrt{x^2 - a^2}} dx$$

Constraining the atmosphere to be in **hydrostatic balance** to obtain a set of equations:

$$\Delta k = F_1(p, T, p_w)$$

$$n' = F_2(p, T, p_w)$$

$$dp/dz = F_3(p, T, p_w)$$



# LMIO principle

The optical depth is given by:

Abel transform to absorption coefficient :

$$\tau = \int_{Tx}^{Rx} k ds \xrightarrow{\text{Amplitude}} T = I/I_0 = \zeta \exp(-\tau) \rightarrow k(z_j) = \frac{1}{\pi} \frac{da}{dr} \Big|_{a=a_j} \int_{a_i}^{a_{top}} \frac{d \ln T(a)}{da} \frac{1}{\sqrt{a^2 - a_j^2}} da$$

The Volume Mixing Ratio (VMR in ppmv) profiles for species can retrieved by:

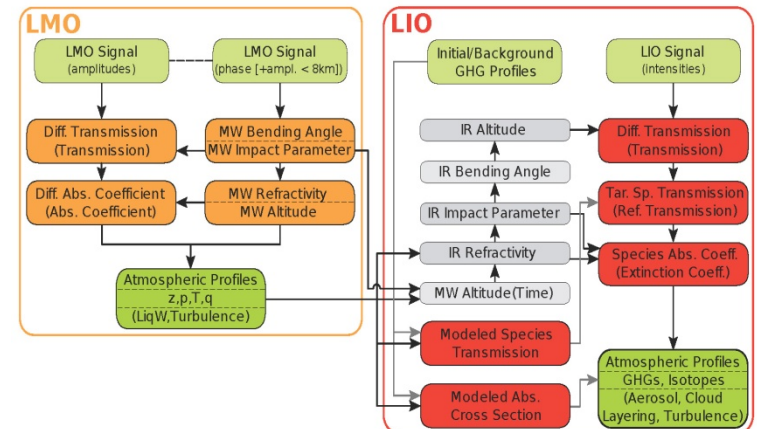
$$\chi(z_i) = 10^6 R \frac{k(z_i) T(z_i)}{\varepsilon(z_i) p(z_i)}$$

Doppler effected optical depth difference is give by:

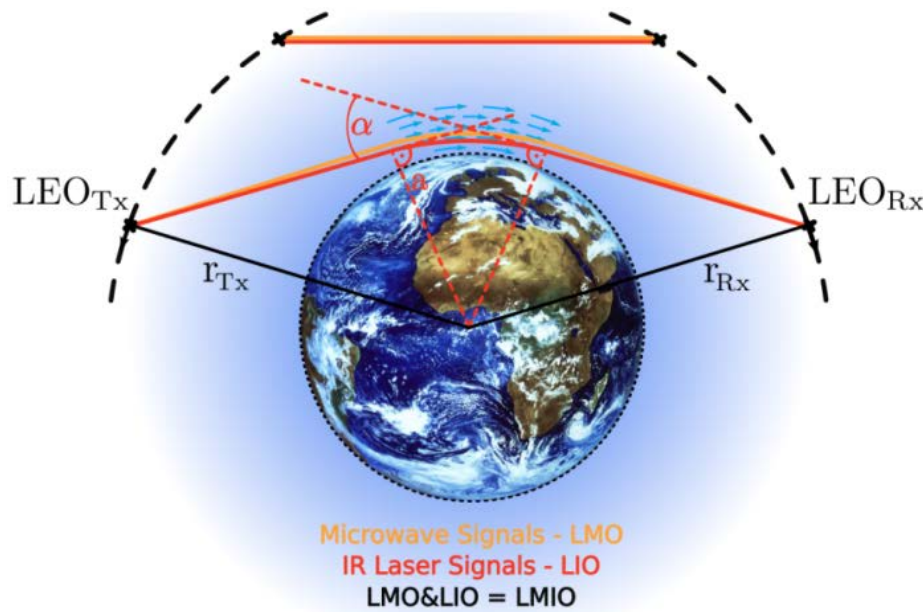
$$\Delta \tau = \int_{Tx}^{Rx} \Delta k_0 ds - \frac{1}{c} \int_{Tx}^{Rx} \Delta \chi_0 v_{||} ds$$

Abel transform to wind velocity :

$$v(a) \approx \frac{c}{\Delta \chi_0(a)} \left[ \frac{1}{\pi} \frac{d}{dr} \int_a^{\infty} \frac{\Delta \tau(x) dx}{\sqrt{x^2 - a^2}} + \Delta k_0(a) \right]$$



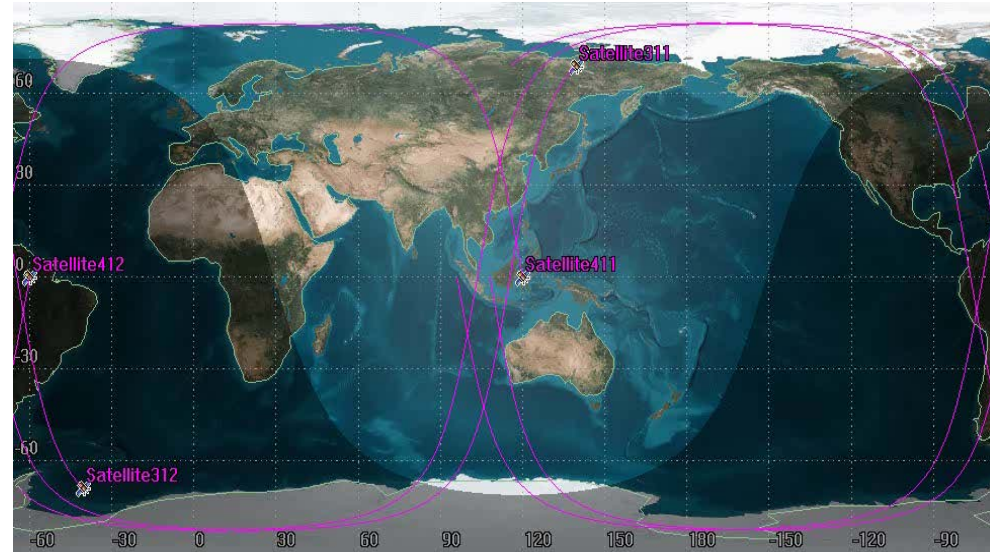
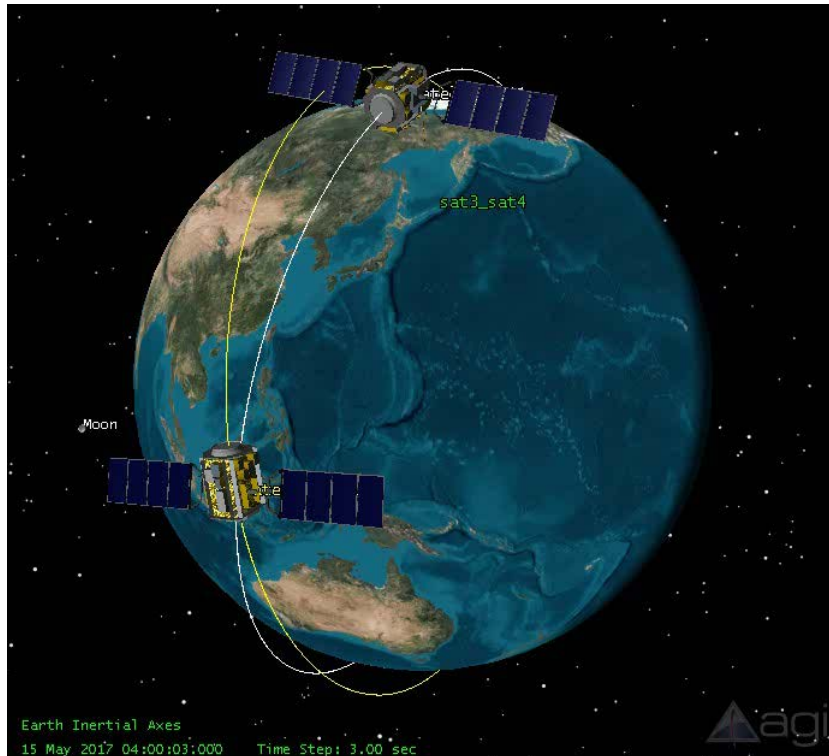
## Climate and Atmospheric Composition Exploring Satellites (CACES) accepted by Strategic Priority Research Program of Chinese Academy of Sciences (SPRPCAS)



- 4 LEO Satellites
- Combining microwave and infrared laser occultation and clouds & water vapor VIR dual-channels imaging
- **LMO and LIO limb sounding** for vertical profiles of T, P, WV, CO<sub>2</sub>, CH<sub>4</sub> (etc.) at high vertical resolution
- **CWAVI camera** for down-looking multi-angle images over occ. event regions

- To monitor the multiple parameters simultaneously in high accuracy, high vertical resolution, to provide a new understanding of spatial-temporal distribution and variation.
- To know the transport and exchange of atmospheric water vapor, liquid water, maybe including ice crystals, in clouds under strong convections.
- To reveal the physical and chemical processes of different atmospheric composition at the key interface between troposphere and stratosphere, and to study their impacts and feedbacks on global climate change.
- To study the impact of atmospheric composition change on the Earth's ecosystem and the interaction mechanisms

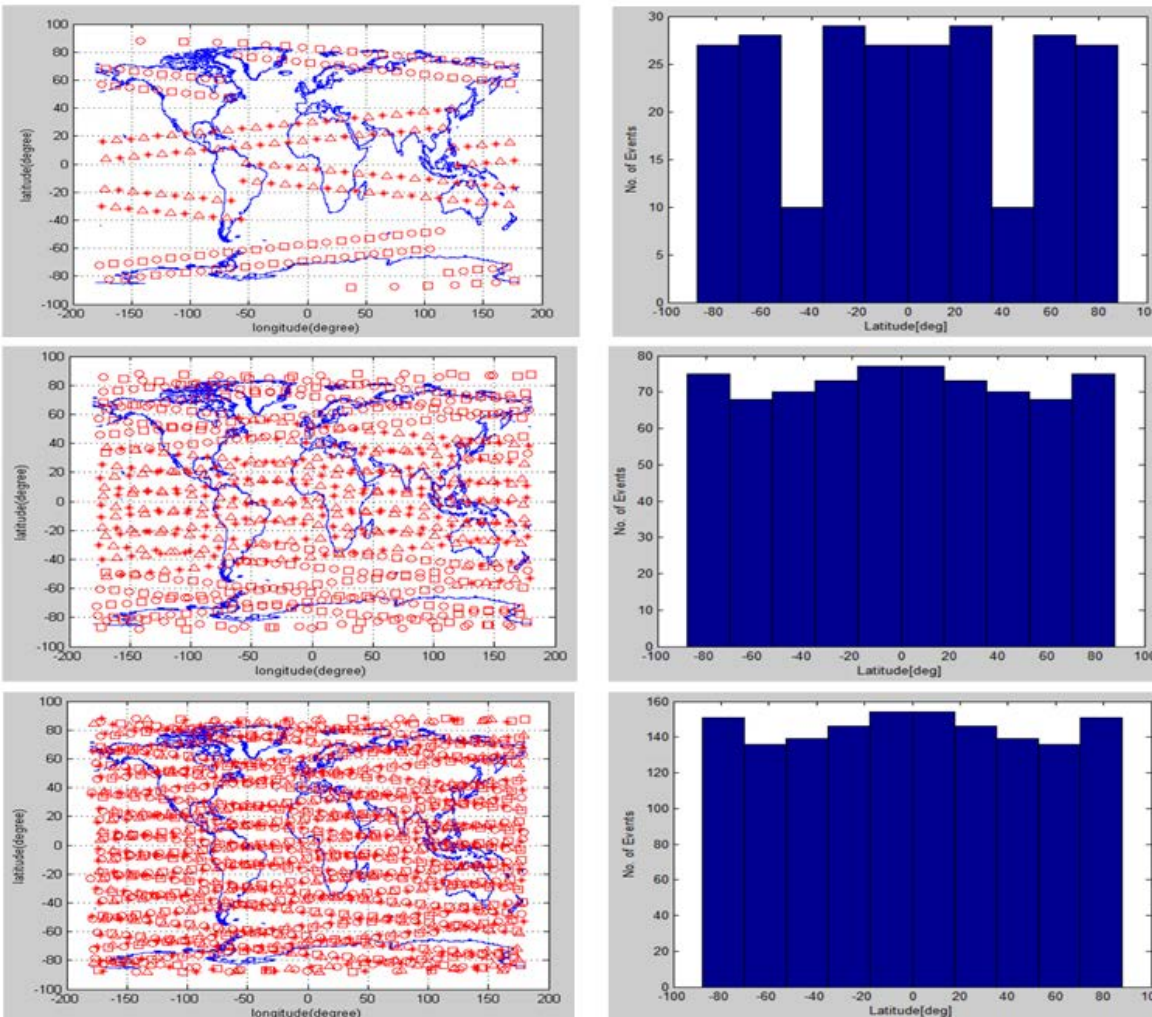
## Constellation Design



- Sun synchronous/near polar counter-rotating orbits of Tx and Rx
- 550/500 km
- 2 Satellites/orbit,  
4 Satellites in total



# LEO-LEO occultation analysis



Orbits	2 Rx Sat	2 Tx Sat
<b>Height</b>	500km	550km
<b>Eccentricity</b>	0	0
<b>Inclination</b>	97.41°	97.60°
<b>Time</b>	12:00	0:00
<b>Period</b>	5676.98 sec	5738.99 sec

Global distribution of the LEO-LEO events for 24/72/144 hours



# Parameters of CACES system



System		Mass (Kg)	Avg. P (W)	Peak P (W)
<b>Platform</b>	Att.	30.00	40.00	76.0
	Energy	28.00	10.00	10.0
	Task	8.00	12.00	12.0
	M&C	6.00	23.00	23.0
	Data Tr.	8.00	6.00	60.0
	Struc.	80.00	0.00	0.0
	Thermal	7.00	30.00	40.0
	Line	9.00	/	/
	Orbit	15.00	/	/
	Propel	15.00	/	10.0
<b>Payloads</b>	LIO (Tx/Rx)	60/40	5/4	100/30
	LMO (Tx/Rx)	40/30	5/4	50/30
<b>Mass in total ( Tx/Rx )</b>		306/276		
<b>Power in total ( Tx/Rx )</b>			131/129	381/291

# Observations of CACES



- Vertical range: Earth surface – 40 to 80 km
- Vertical Resolution: better than 500 m
- T, P, WV, CO<sub>2</sub>, CH<sub>4</sub>, (etc.)

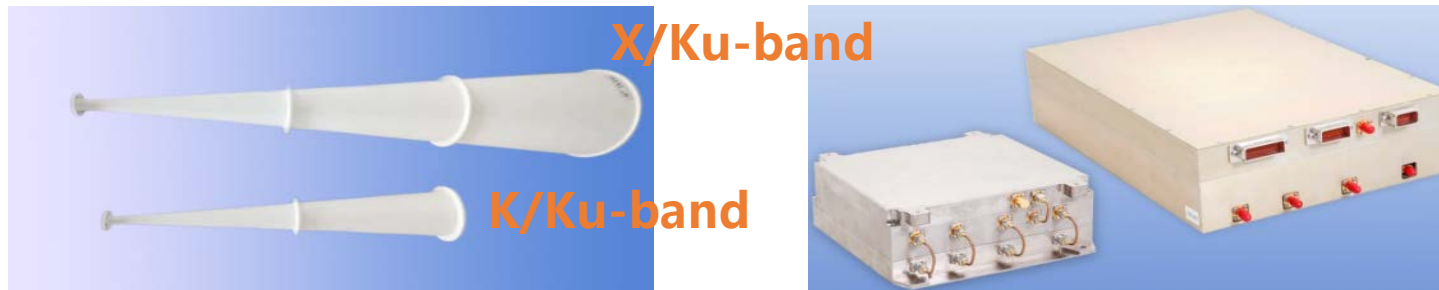
		MO				IO				Units
		Temperature		Sp. Humidity		Trace Species		I.o.s. Wind		
Requirement		Target	Thres	Target	Thres	Target	Thres	Target	Thres	
Vertical domain		S-80	S-40	S-15	S-10	5-40	5-30	5-40	5-30	[km]
Vertical sampling	LT	0.5	1	0.5	1	0.5	1	0.5	1	[km]
	UT	0.5	1	0.5	1	0.5	1	0.5	1	[km]
	LS	0.5	1	0.5	1	0.5	1	0.5	1	[km]
	US	1	2	-	-	1	2	1	2	[km]
RMS accuracy	LT	best-effort basis								Temp [K] Humi [%] Species [%] Wind [m/s]
	UT-bottom									
	UT- ≥10km	1	2	10	20	TBD				
	LS	0.5	1	10	-					
	US	0.5	1	10	-					
	US	1.5	3	-	-					

To propose a LMO mission in China, NSSC has conducted a pre-study in four aspects:

- Orbit design and GRO&LMO events analysis
- Frequencies selection and error analysis
- Transmitter and receiver design&development
- Preparing an outdoor demonstration experiment

(The simulation work was conducted by using EGOPS (WEGC))

# LMO payloads design

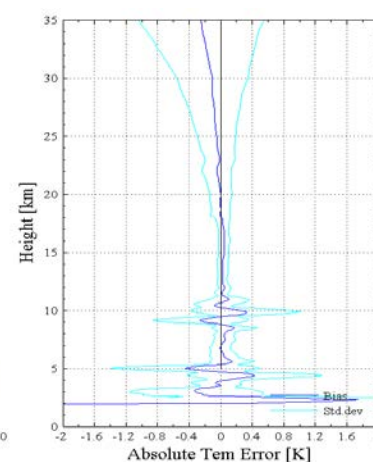
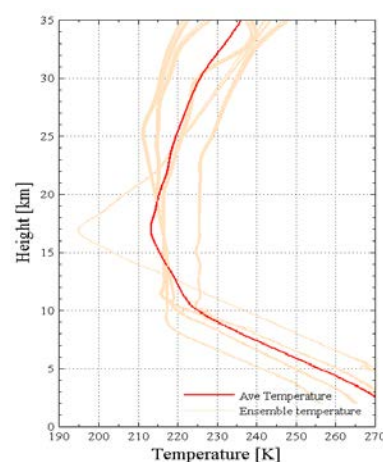
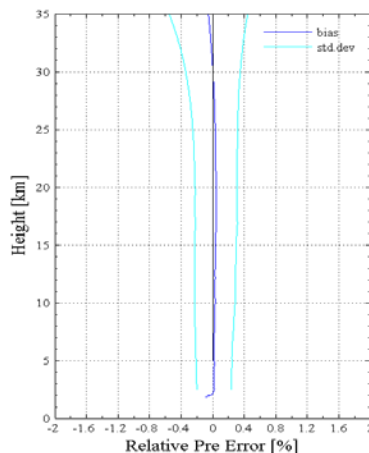
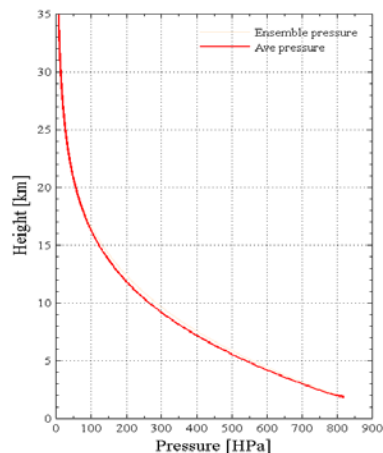


<b>Frequencies</b>	~9.4GHz	~14.6GHz	~17.4GHz	~22.6GHz
<b>Antenna Gain</b>	≥26dB	≥27.5dB	≥28.5dB	≥30dB
<b>Transmit Power</b>	≥2W	≥2W	≥2W	≥2W
<b>Amplitude Measurement Accuracy</b>	Less than 0.022dB ( 0.5% ) RMS ( within 15 seconds )			
<b>Phase Measurement Accuracy</b>	Better than 1% cycle			
<b>Clock Short-term Frequency Stability</b>	Better than 1e-12/ second			
<b>Sampling Rate</b>	≥500Hz			
<b>Weight</b>	30kg (T & R)			
<b>Power consumption</b>	60W (T) 40W (R)			
<b>Size</b>	Antenna: X/Ku : φ270*920mm Ku/K: φ200* 900mm			

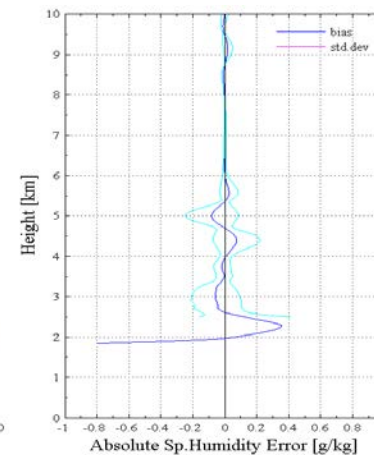
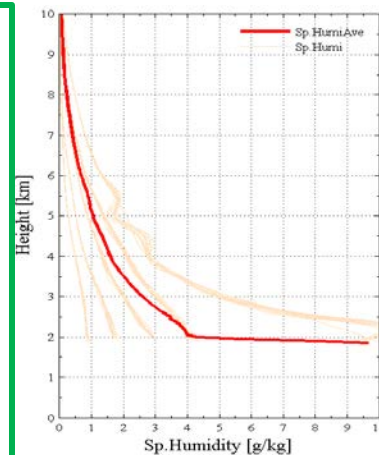
# Frequencies and analysis



Frequencies	~9.7GHz	~13.5GHz	~17.25GHz	~22.6GHz
Gain	≥26dB	≥27.5dB	≥28.5dB	≥30dB
Power	≥2W	≥2W	≥2W	≥2W
Sampling	≥500Hz			

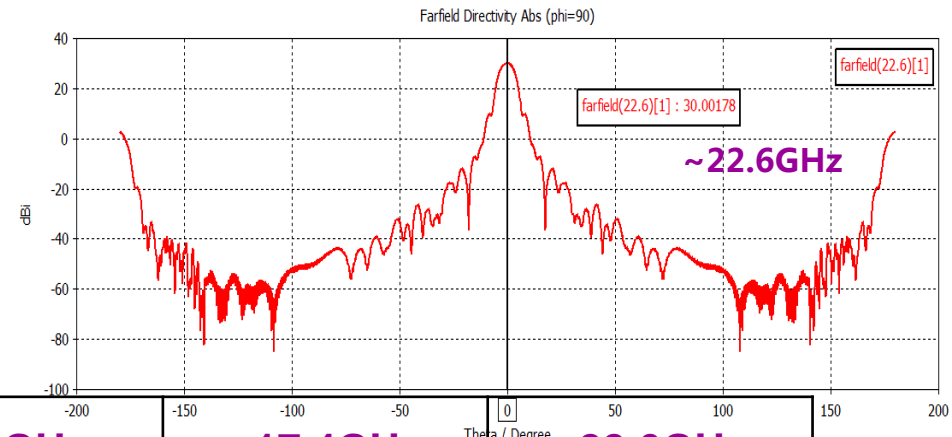
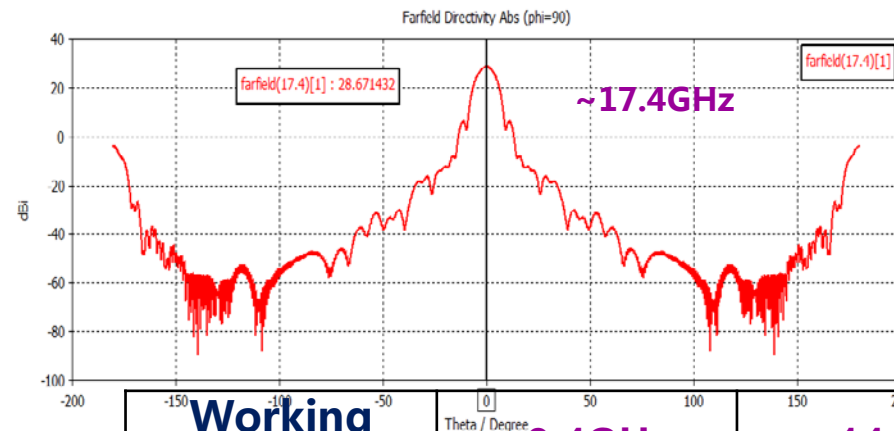
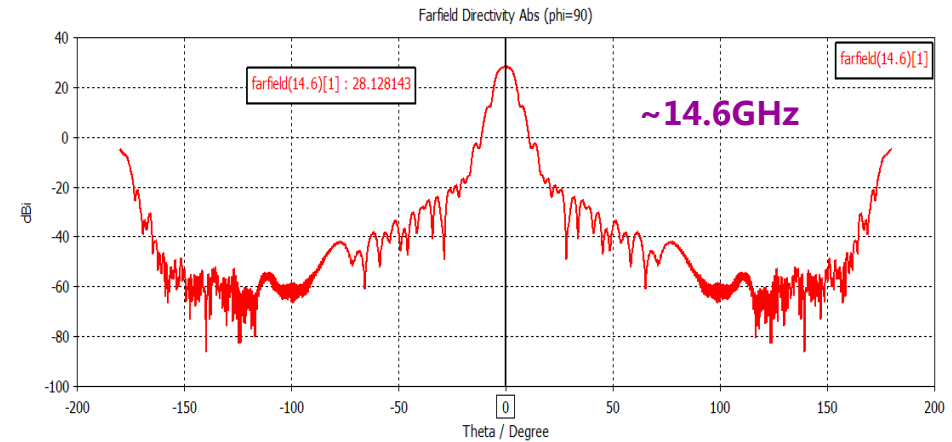
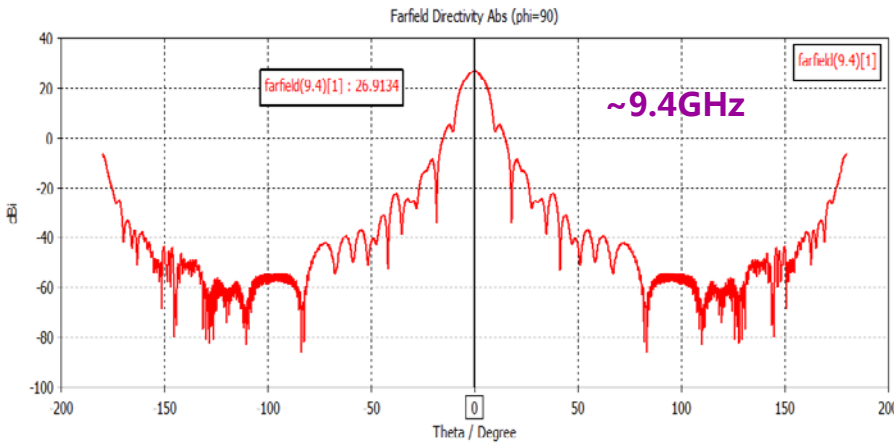


In order to probe the lower troposphere **using differential measurements, four frequency channels** (see the table) were selected for the Tx and Rx design. Using these four frequencies, an error effects analysis has been conducted using End-to-End simulation (EGOPS). Each error (e.g. POD uncertainty, thermal noise, gain drift, clock bias... ) effect has been analyzed, individually. Then the total errors' effect on the LMO retrievals was performed. Here we show the total errors' effects on **pressure, temperature and humidity**, by RO profiles.





# Antenna amplitude stability control



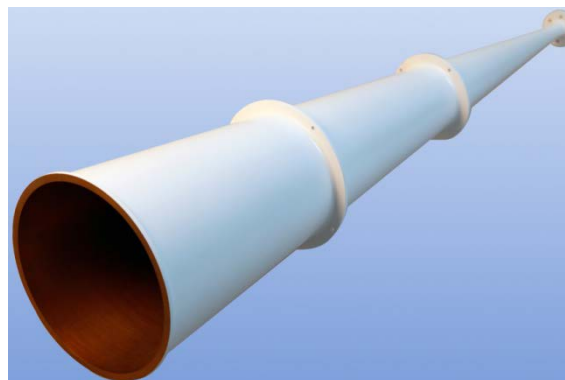
Working Frequency	~9.4GHz	~14.6GHz	~17.4GHz	~22.6GHz
Gain Requirement	≥26dB	≥27.5dB	≥28.5dB	≥30dB
Simulation Results	26.9 dB	28.1 dB	28.7 dB	30 dB

# Tx, Rx and their indoor test



## Antenna performance

Parameters	Gain (dBi)	Targets (dBi)
9.4G-v	27.03	≥26
14.6G-v	29.11	≥27.5
17.4G-v	28.61	≥28.5
22.6G-v	30.90	≥30
9.4G-vf	26.95	≥26
14.6G-vf	29.39	≥27.5
17.4G-vf	28.71	≥28.5
22.6G-vf	30.85	≥30
9.4G-h	26.49	≥26
14.6G-h	29.08	≥27.5
17.4G-h	28.73	≥28.5
22.6G-h	30.75	≥30
9.4G-hf	26.52	≥26
14.6G-hf	29.16	≥27.5
17.4G-hf	28.83	≥28.5
22.6G-hf	30.79	≥30



Frequencies		9.4GHz	14.6GHz	17.4GHz	22.6GHz
Powers	Targets	≥2W	≥2W	≥2W	≥2W
	Test values	2.29W	2.63W	2.45W	2.63W

Frequencies		9.4GHz	14.6GHz	17.4GHz	22.6GHz
Amplitude Stability	Targets	≤0.01dB	≤0.01dB	≤0.01dB	≤0.01dB
	Test values	0.0025dB	0.0027dB	0.0053dB	0.0025dB

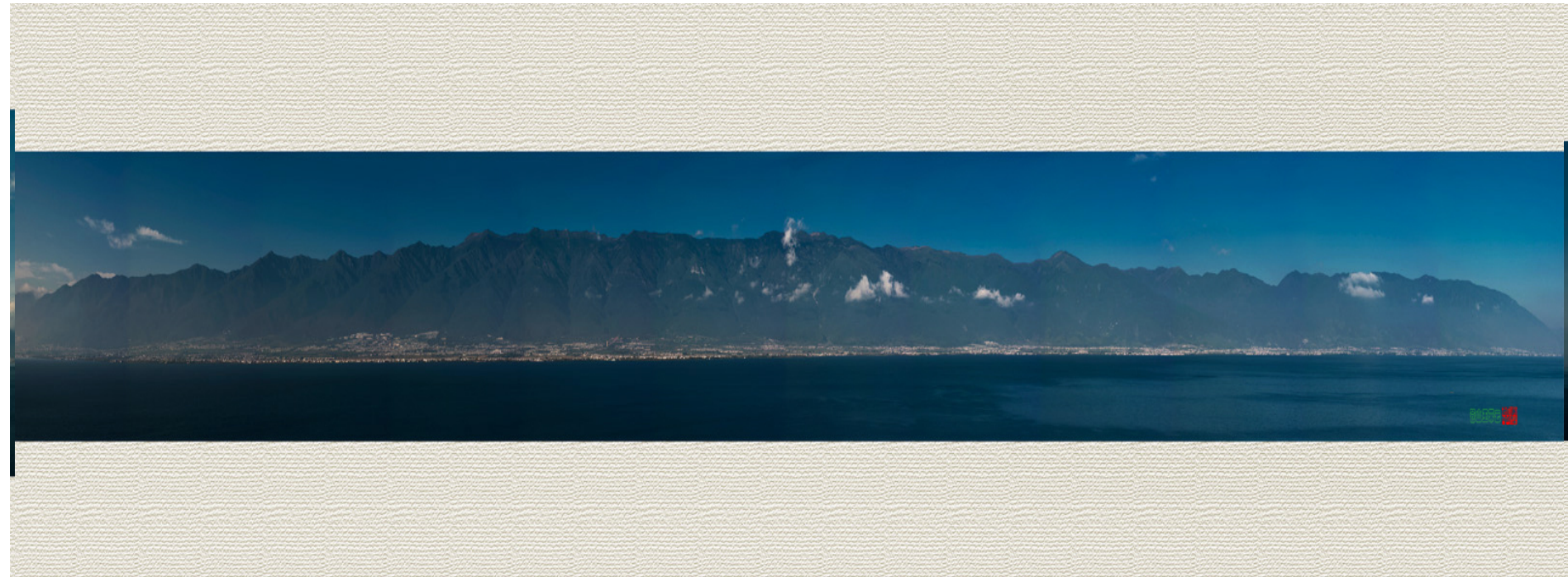
**Transmitter performance**

Frequencies		9.4GHz	14.6GHz	17.4GHz	22.6GHz
Gain Stability	Targets	≤0.01dB	≤0.01dB	≤0.01dB	≤0.01dB
	Test values	0.0027dB	0.0025dB	0.0033dB	0.0032dB



**Receiver performance**

## Dali, Yunnan province



## In China:

- A flight model Sat-Sat occultation receiver has been developed for Mars
- An Earth observation mission named Climate and Atmospheric Composition Exploring Satellites (CACES) using LEO-LEO occultation and clouds & water vapor imaging has been accepted by the Strategic Priority Research Program of Chinese Academy of Sciences (SPRPCAS) as a concept study.
- The study of Earth LMO has been conducted in four aspects:
  - Orbit design and GRO&LMO events analysis;
  - Frequencies selection and error analysis;
  - Transmitter and receiver design&development;
  - Preparing an outdoor demonstration experiment.
- Preliminary studies of LIO have been started as well.

**Look forwards your cooperation...**

**Thank you!**



# Conferences announcement:



Conference: **IEEE GNSS+R 2021**

Country: China

City: Beijing

Start Date: **10<sup>th</sup> May 2021**

End Date: **12<sup>th</sup> May 2021**

Conference: **IGL-2 2021**

Country: China

City: Beijing

Start Date: **13<sup>th</sup> May 2021**

End Date: **14<sup>th</sup> May 2021**

We would like to provide conference venues, service and another contributions to the **following on IROWG workshops** as well.



Welcome  
to Beijing!



*or other weeks in May (TBC)*